

*H atom cluster formation and
recombination on the graphite
(0001) surface*

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University of Aarhus, Denmark

Eagle-Nebula

10-1000 K, 100-10.000 atoms/cm³



Star-Birth Clouds · M16

HST · WFPC2

PRC95-44b · ST ScI OPO · November 2, 1995
J. Hester and P. Scowen (AZ State Univ.), NASA

Cloud composition

Atoms:

H, He, O, C, N, Ne, Si, Mg, S, Fe ...

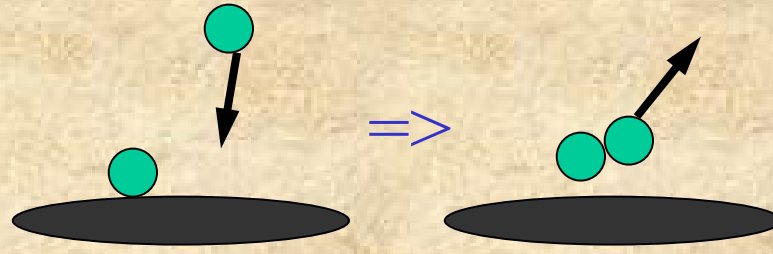
~130 Molecules:

H₂, CO, H₂O, CO₂, O₂, NH₃, CH₃OH ...

Sugars: glycolaldehyde (CH₂OHCHO)

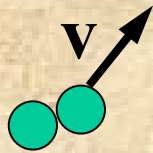
Dust grains

Energy branching in H₂ formation ?

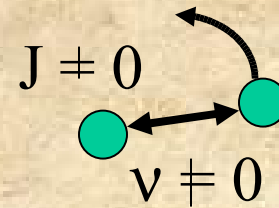


$$E_{\text{released}} \sim 4.5 \text{ eV}$$

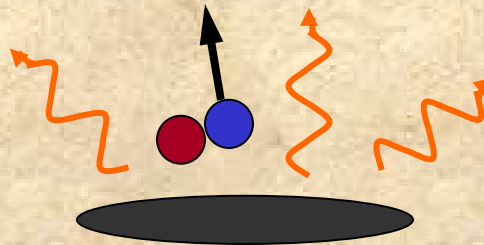
Into:



Kinetic energy ?

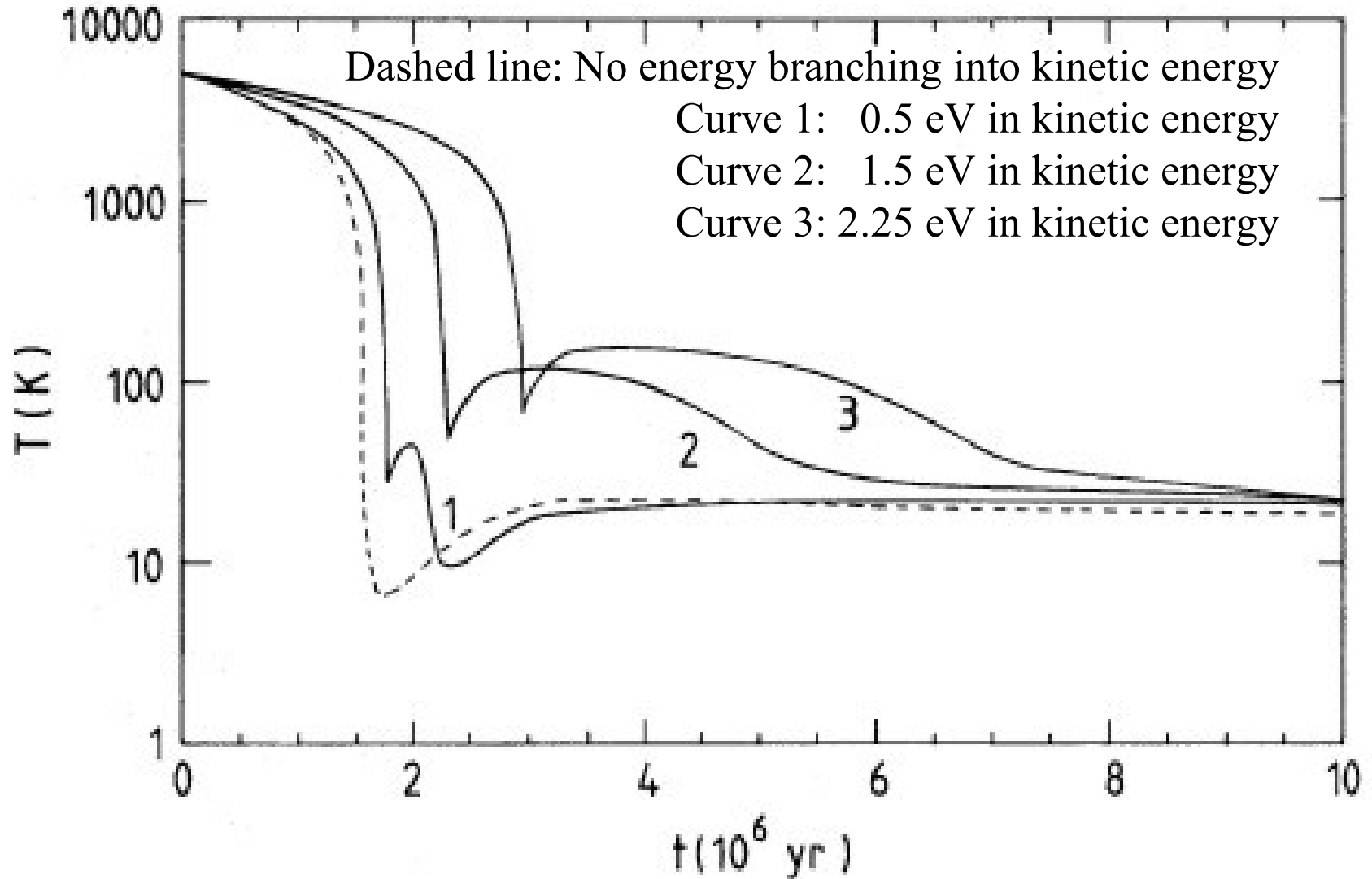


Molecular excitation ?

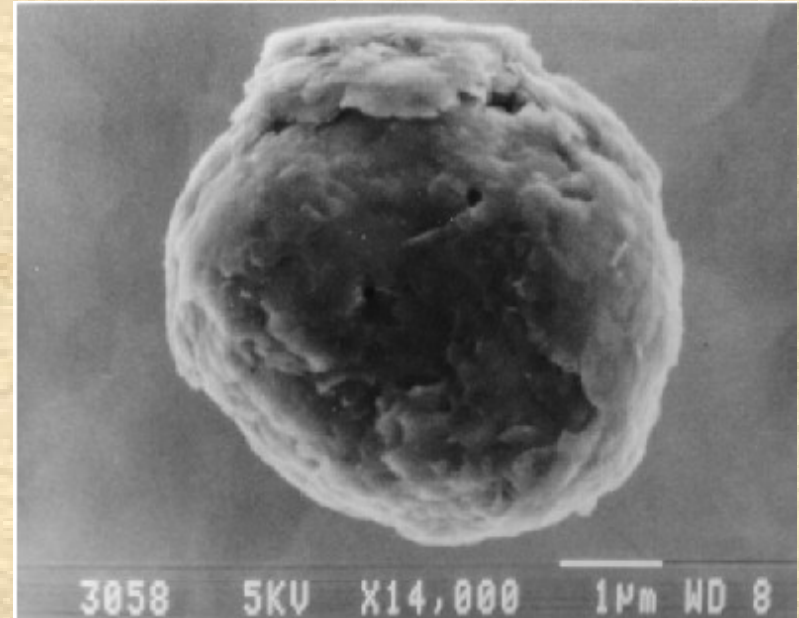
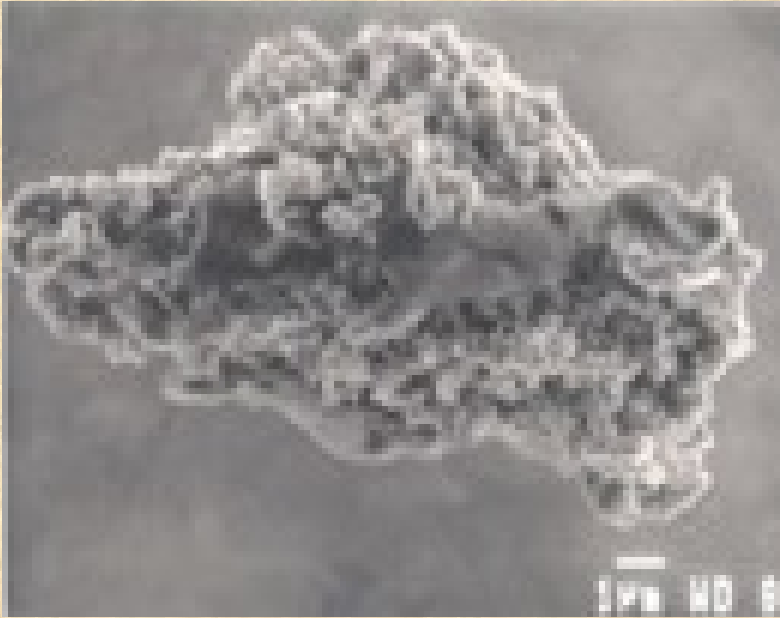


Grain heating ?

Energy release in H₂ formation and the thermal evolution of interstellar clouds



Dust grains



Carbonaceous grains: Graphite, Amorphous Carbon, HAC, PAH, Polymeric Carbon, Diamond

Silicates:

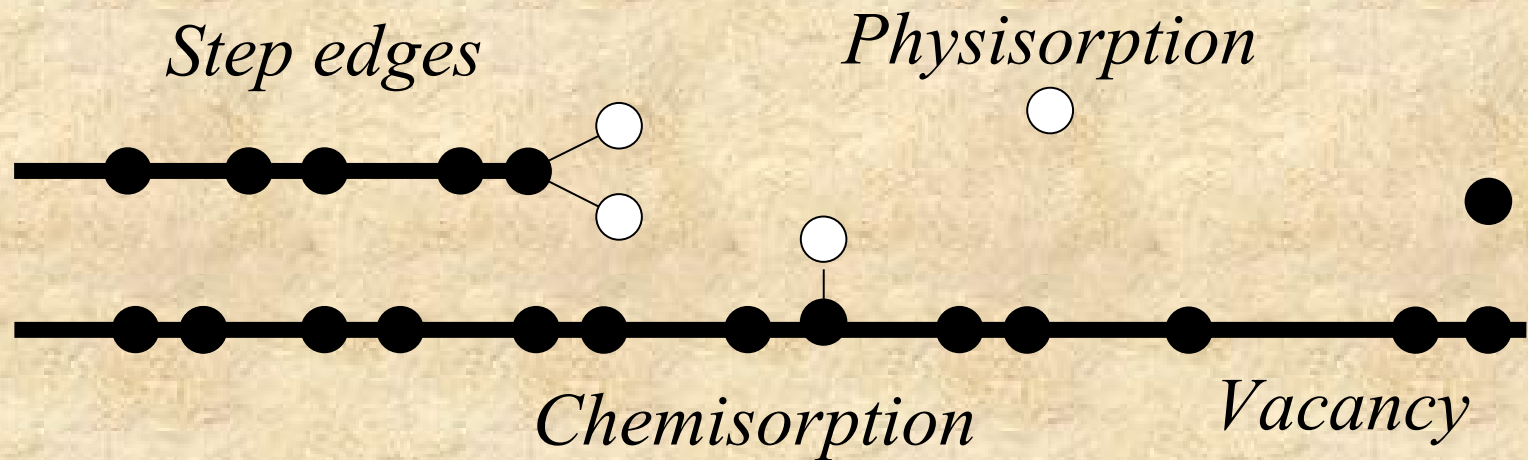
Olivines (Mg_2SiO_4 ,

Fe_2SiO_4)

Ices:

H_2O , CO , CO_2 , CH_3OH , CH_4 , H_2CO ...

Binding sites on graphitic surfaces



Physisorption: Creighan et al, J. Chem. Phys. 124, 114701 (2006)

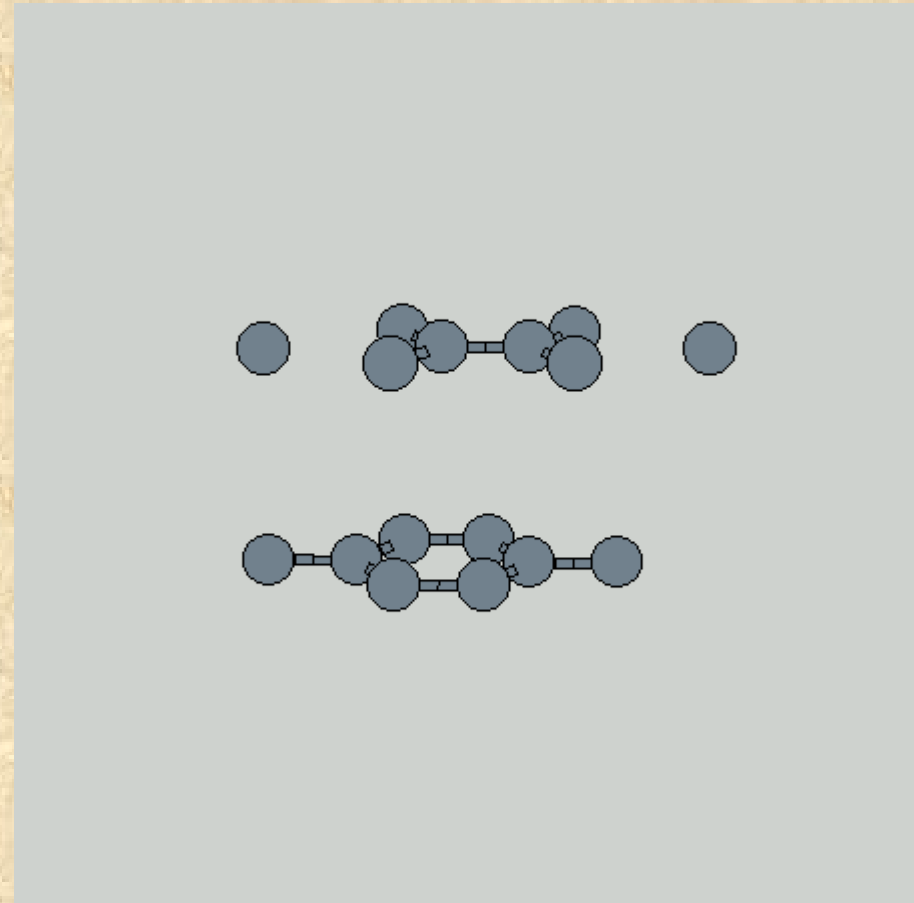
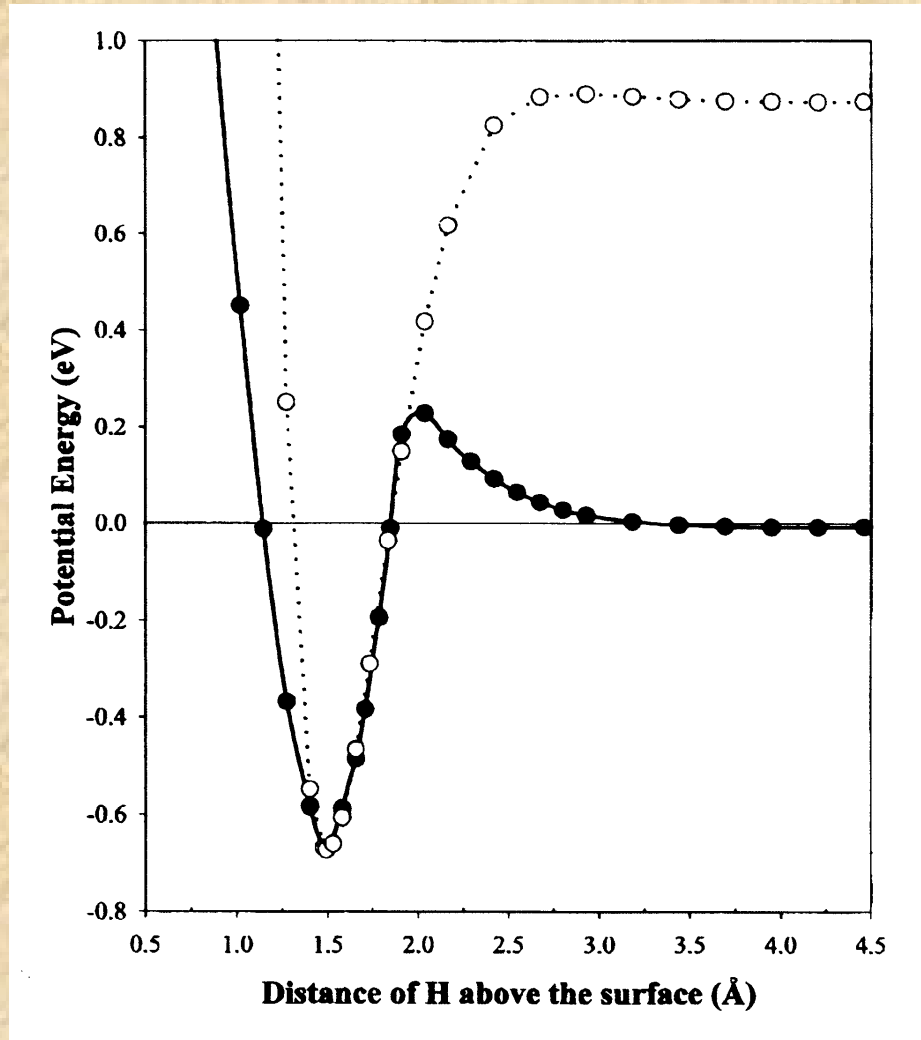
Chemisorption - basal plane: Jeloica & Sidis, Chem. Phys. Lett. 300, 157 (1999)

Chemisorption at defects: Sha et al, J. Am. Chem. Soc. 126, 13095 (2004)

Güttler et al, Surface Science 570, 218 (2004)

H chemisorbed on HOPG

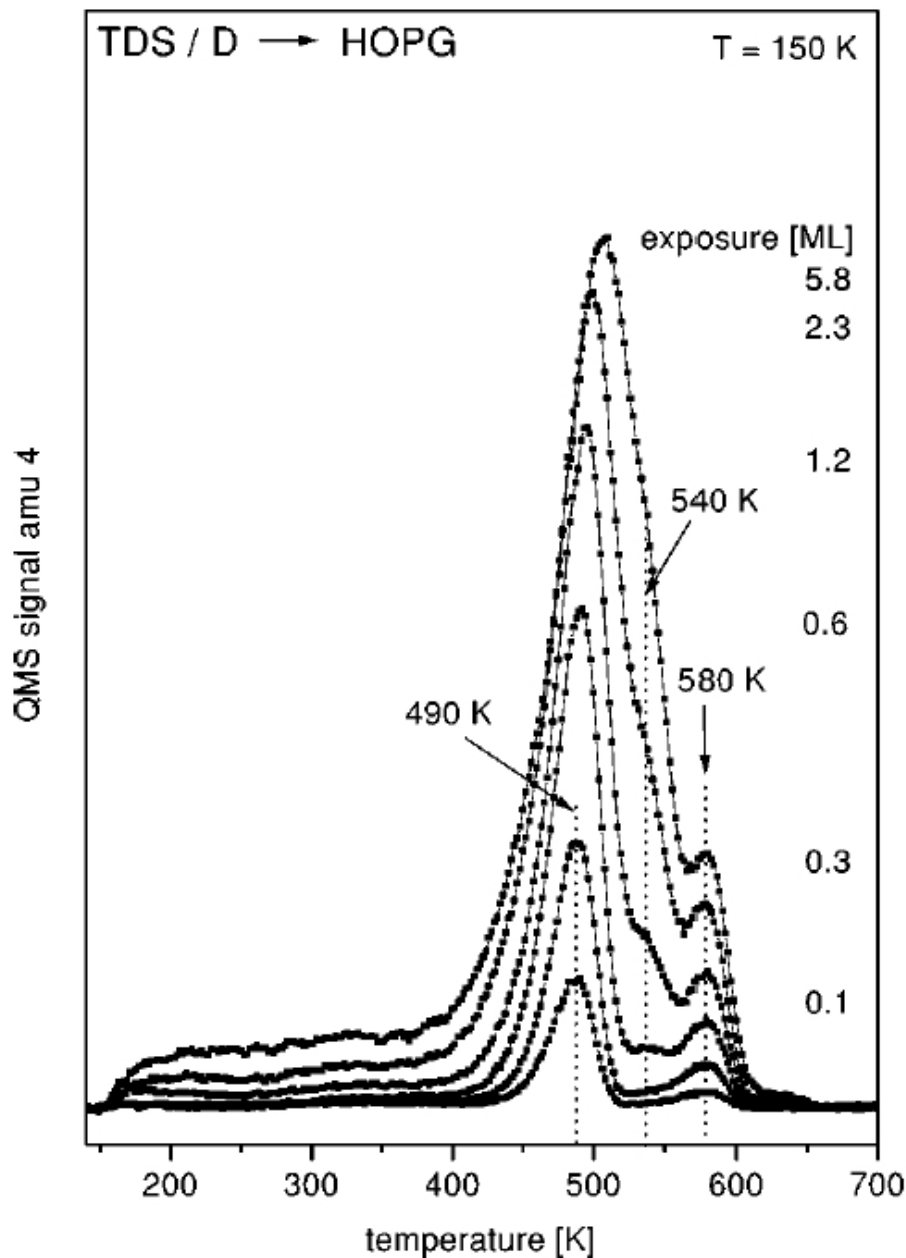
Jeloica & Sidis, Chem. Phys. Lett. 300, 157 (1999)



Eva Rauls

Brett Jackson et al.

H₂ formation on graphite



$$\frac{d\Theta}{dt} = -k_0 e^{-E_B/k_B T} \Theta^n$$

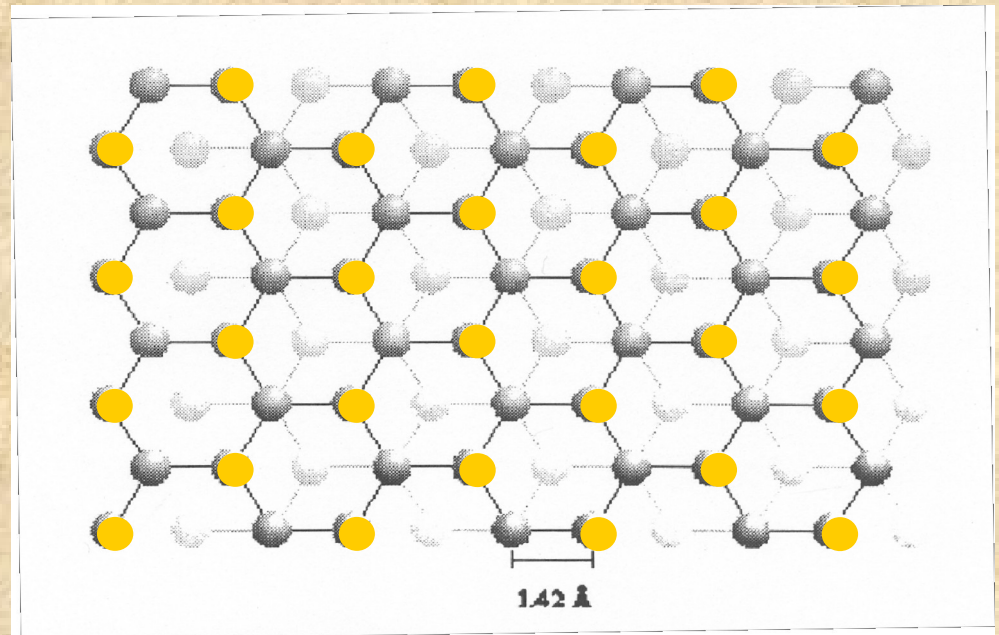
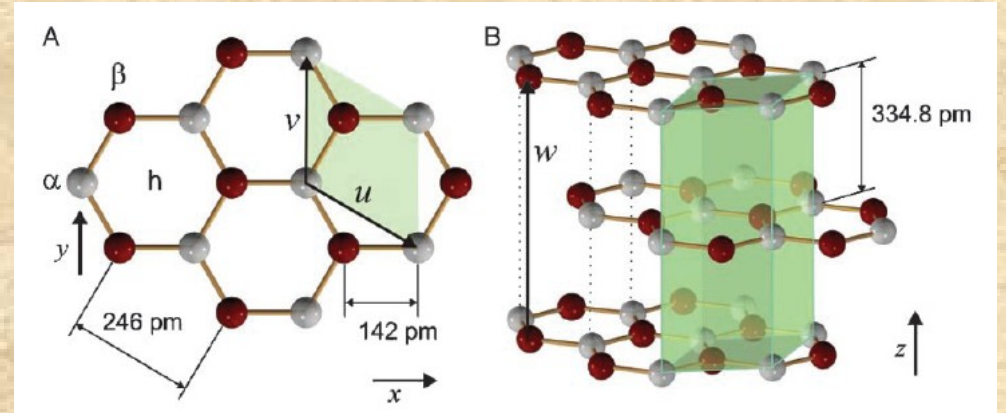
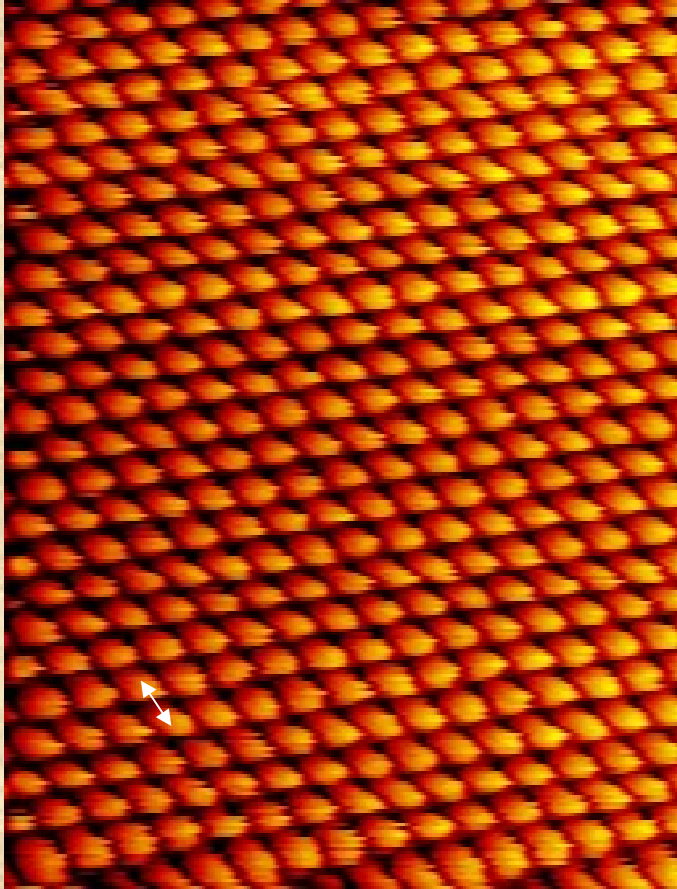
$$n=1$$

⇒ First order desorption

$$490 \text{ K} \Rightarrow 1.4 \text{ eV}$$

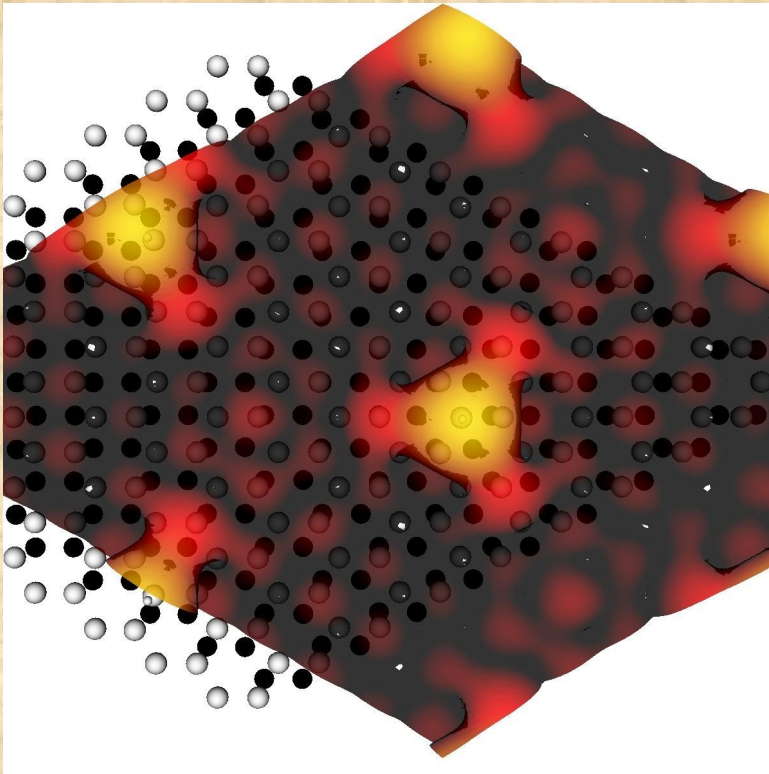
$$580 \text{ K} \Rightarrow 1.6 \text{ eV}$$

STM on graphite

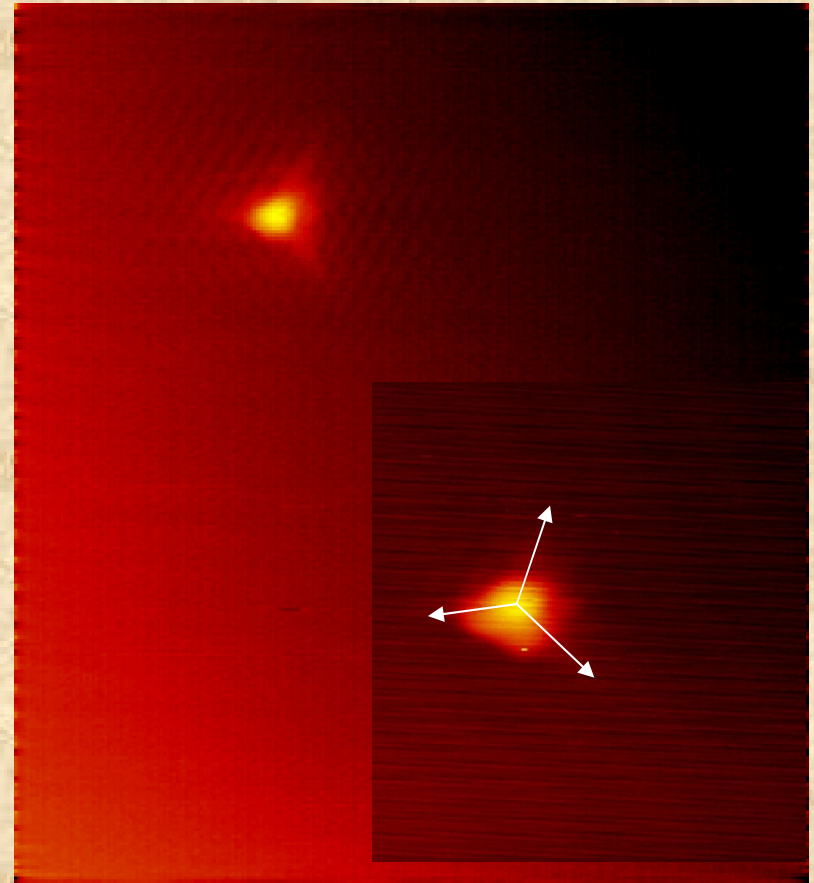


$\updownarrow 2.46 \text{ \AA}$

Hydrogen on graphite – Monomers



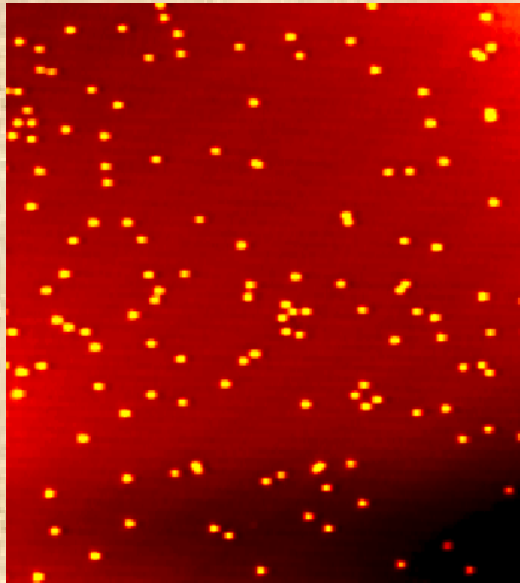
Zeljko Sljivancanin



155 x 171 Å², 180 K
 $V_t \sim -710\text{mV}$, $I_t \sim -0.16\text{nA}$

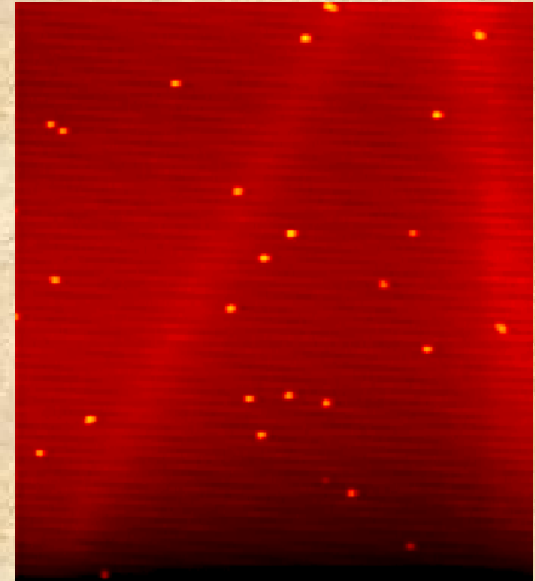
Monomer desorption

Flux:
 $10^{12} \text{ cm}^{-2}\text{s}^{-1}$
 $\Theta \sim 0.03\%$
STM at
 $\sim 180 \text{ K}$



$1030 \times 1140 \text{ \AA}^2$

RT
→
20%



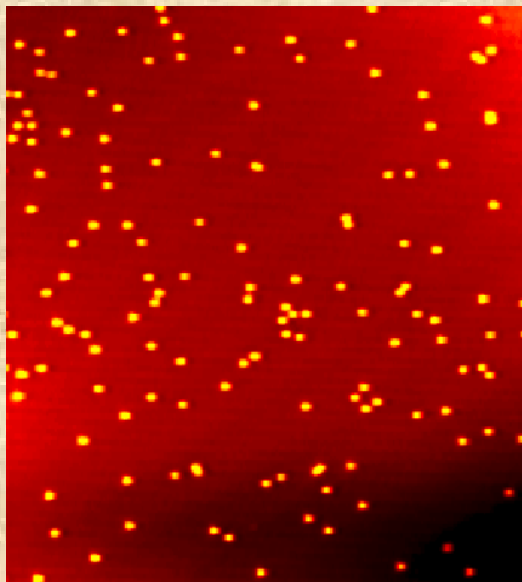
$1030 \times 1140 \text{ \AA}^2$

Experiment: Upper limit: $\tau = 6 \text{ min.}$

Theory: $E_b = 0.9 \text{ eV}$, $\nu = 10^{13} \text{ s}^{-1} \Rightarrow \tau = 130 \text{ s}$

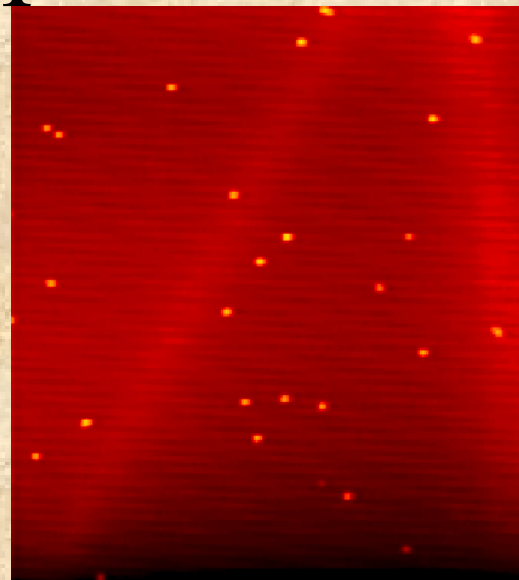
Monomer desorption

Flux:
 $10^{12} \text{ cm}^{-2}\text{s}^{-1}$
 $\Theta \sim 0.03\%$
STM at
 $\sim 180 \text{ K}$



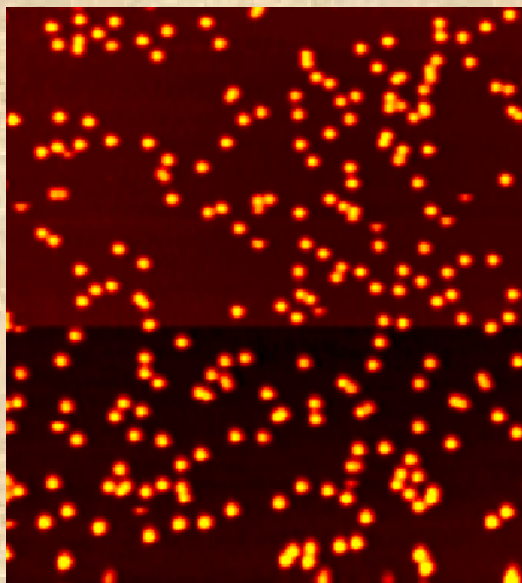
$1030 \times 1140 \text{ \AA}^2$

RT
→
20%



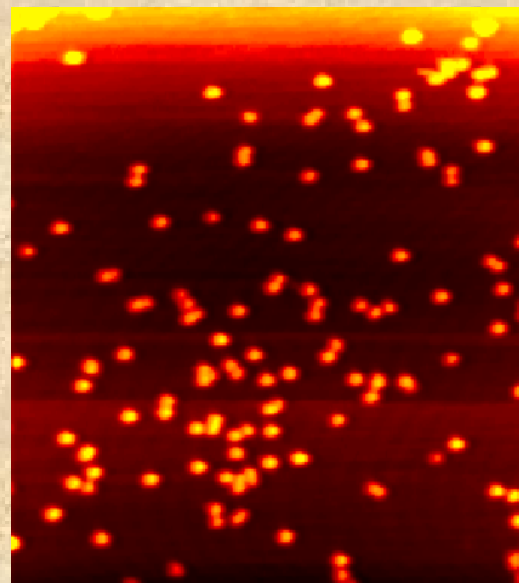
$1030 \times 1140 \text{ \AA}^2$

Flux:
 $10^{14} \text{ cm}^{-2}\text{s}^{-1}$
 $\Theta \sim 0.2\%$
STM at
 $\sim 170 \text{ K}$



$1030 \times 1140 \text{ \AA}^2$

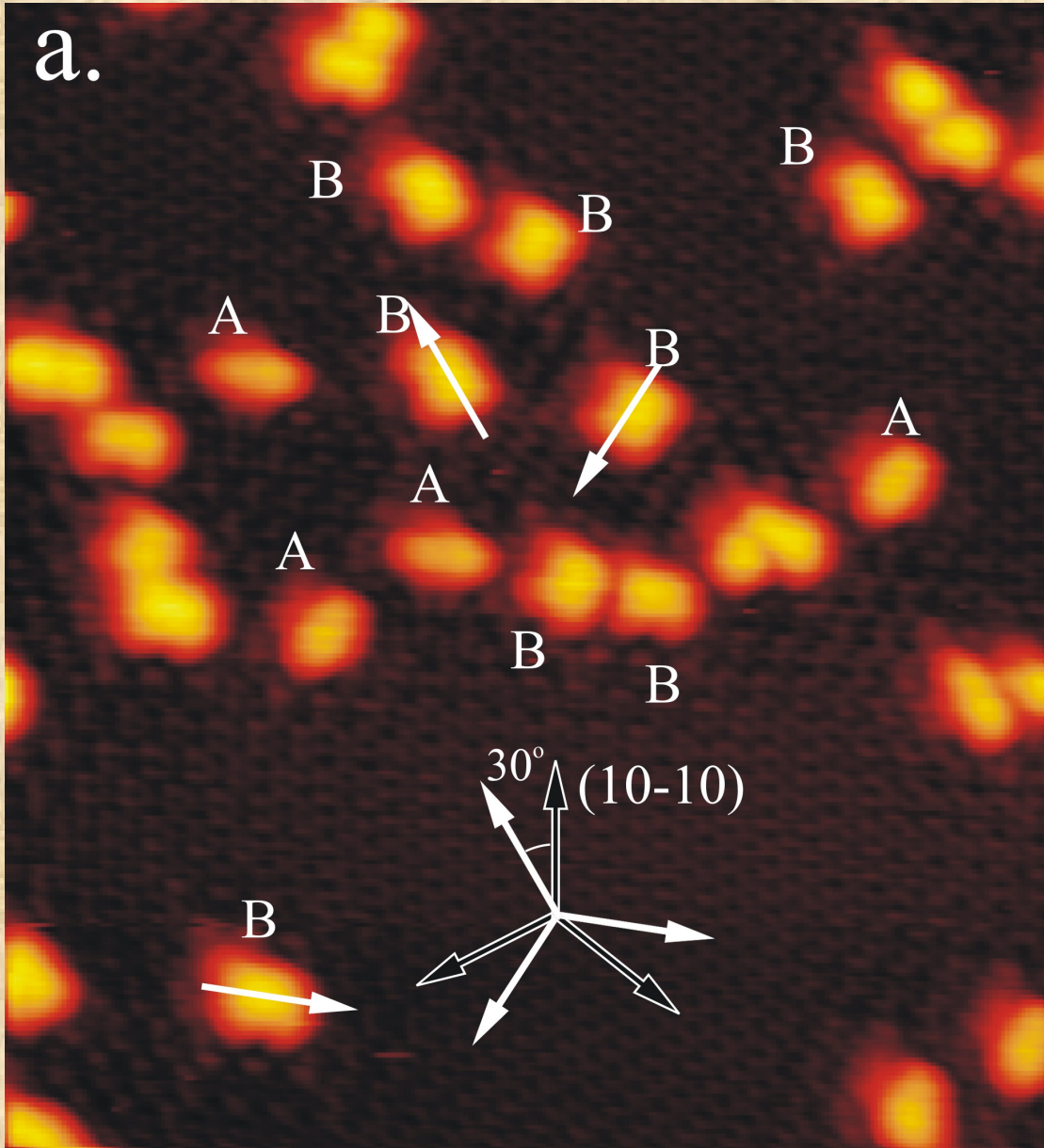
RT
→
80%



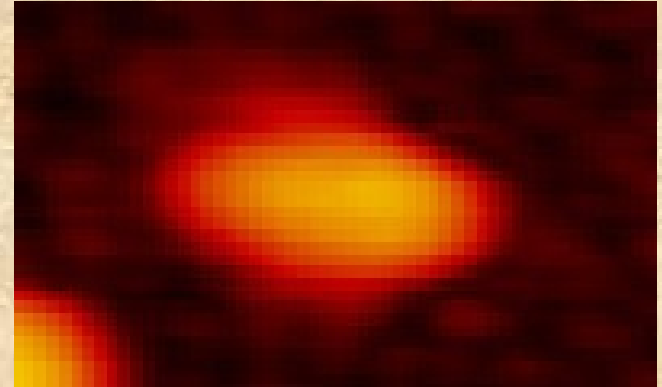
$515 \times 570 \text{ \AA}^2$

H-Dimers on graphite

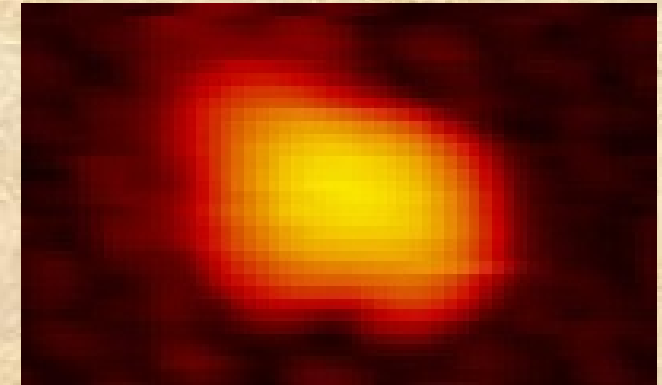
103 x 114 Å²



Dimer A



Dimer B

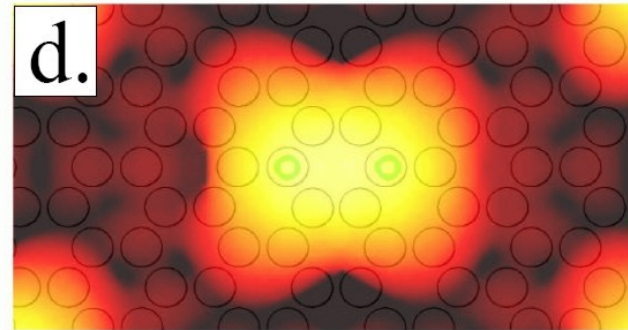
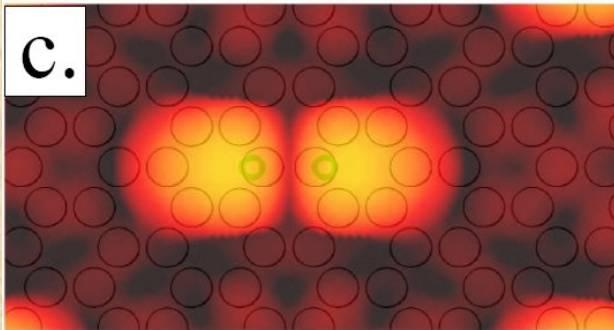
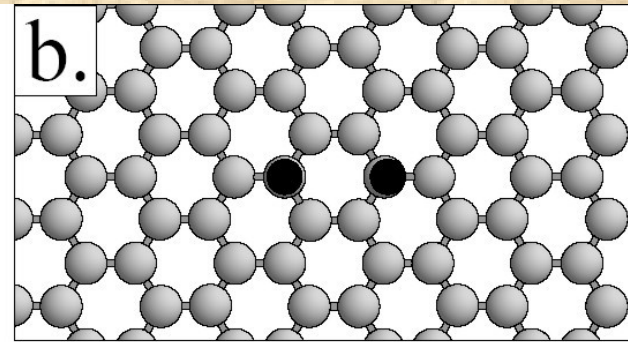
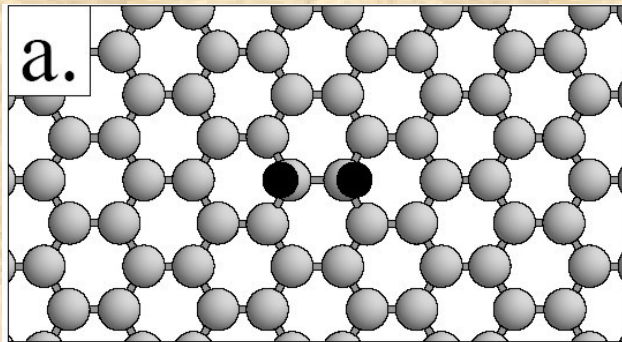


$V_t = 884 \text{ mV}$, $I_t = 0.16 \text{ nA}$

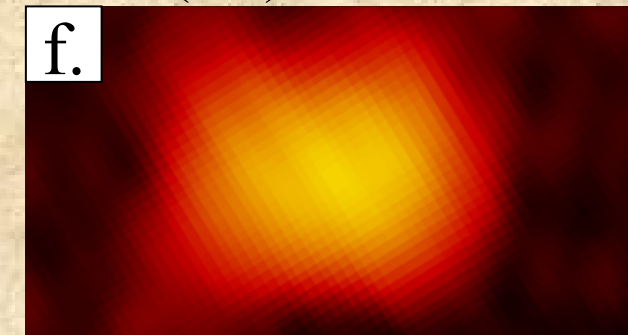
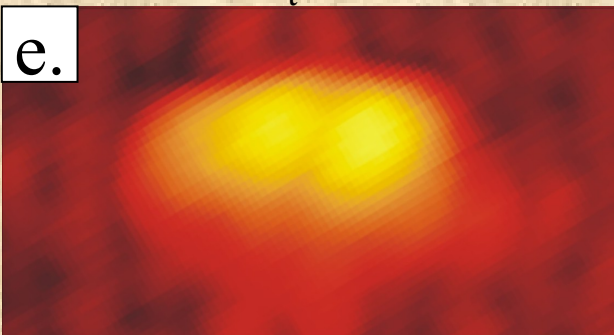
Dimers: Theory vs. Experiment

Ortho dimer - Dimer A

Para dimer - Dimer B

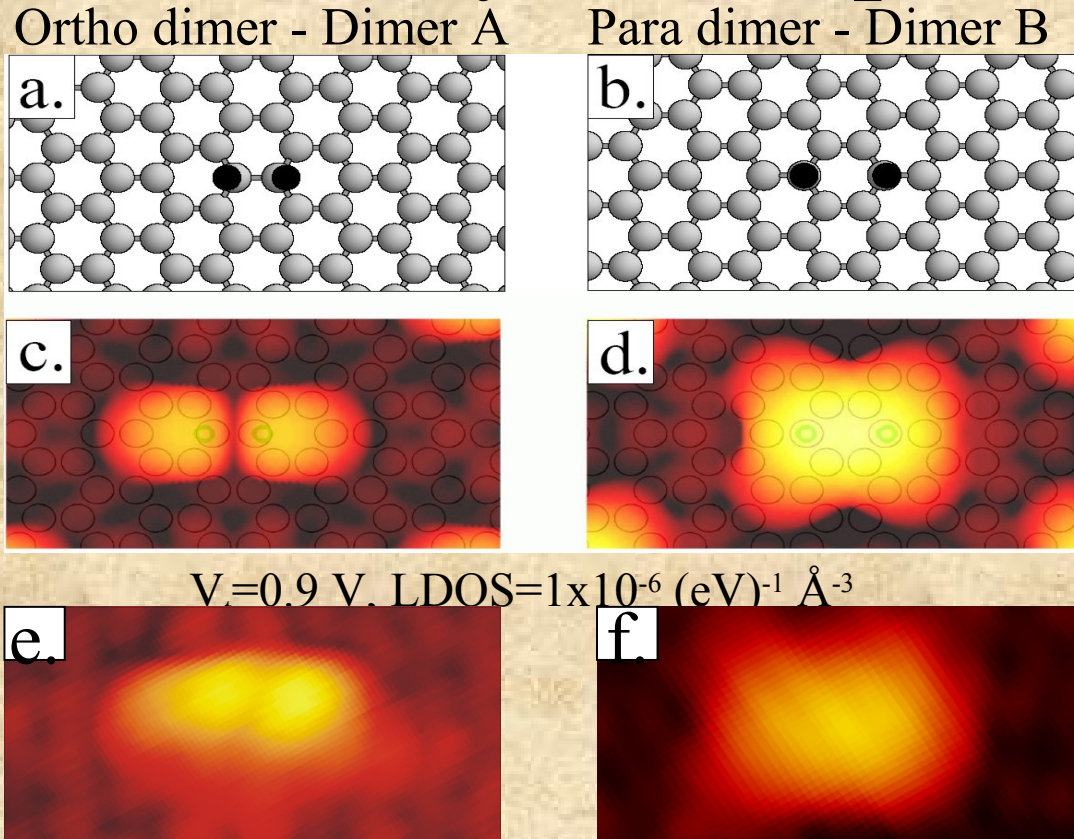


$V_t = 0.9 \text{ V}$, $\text{LDOS} = 1 \times 10^{-6} (\text{eV})^{-1} \text{ \AA}^{-3}$



$V_t = 884 \text{ mV}$, $I_t = 0.16 \text{ nA}$

Dimers: Theory vs. Experiment

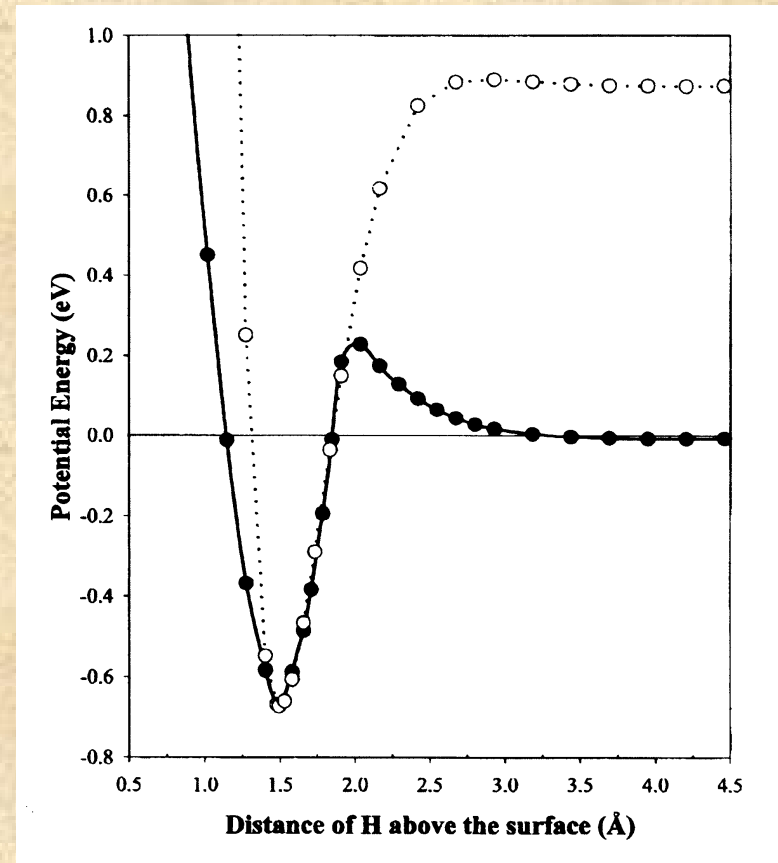
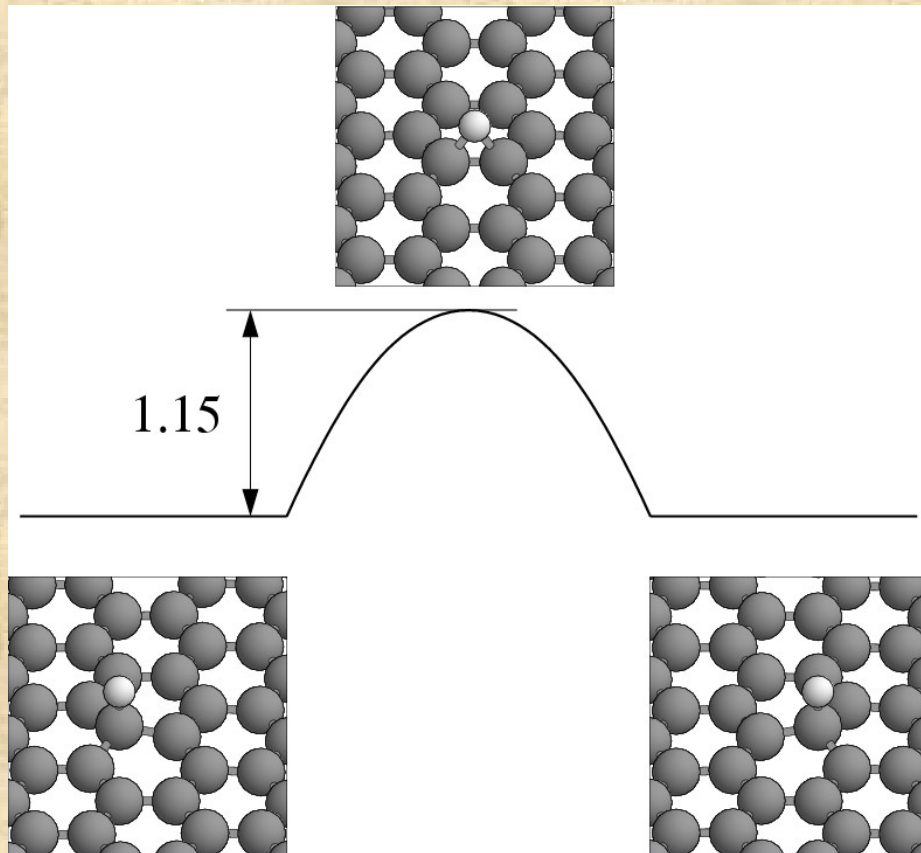


In agreement with: Y. Ferro et al., Chem. Phys. Lett. **368**, 609 (2003).
Partial agreement with: Y. Miura et al. J. Appl. Phys. **93**, 3395 (2003).

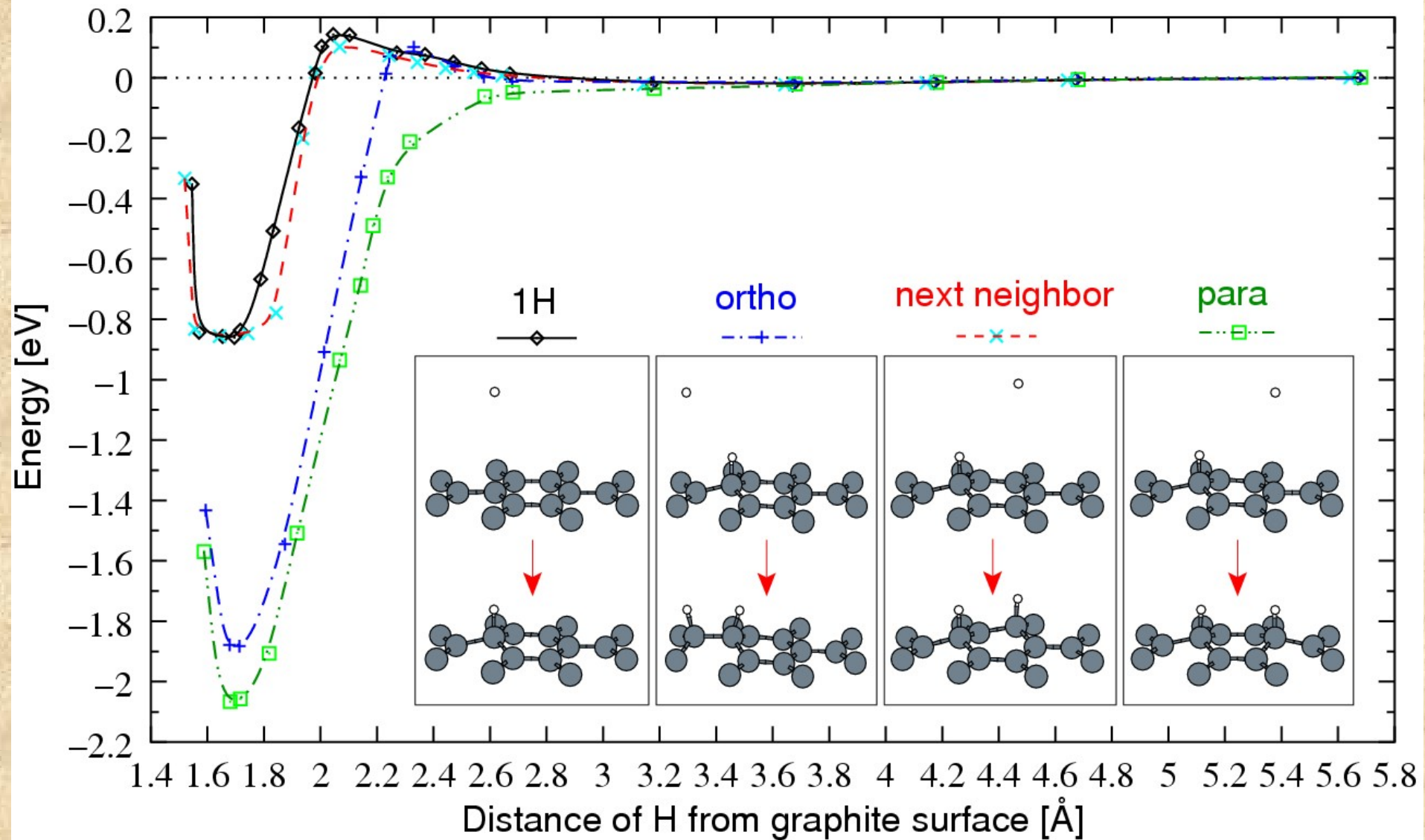
Diffusion

Barrier to diffusion for an isolated H atom: 1.15 eV

Barrier to desorption for an isolated H atom: 0.9 eV

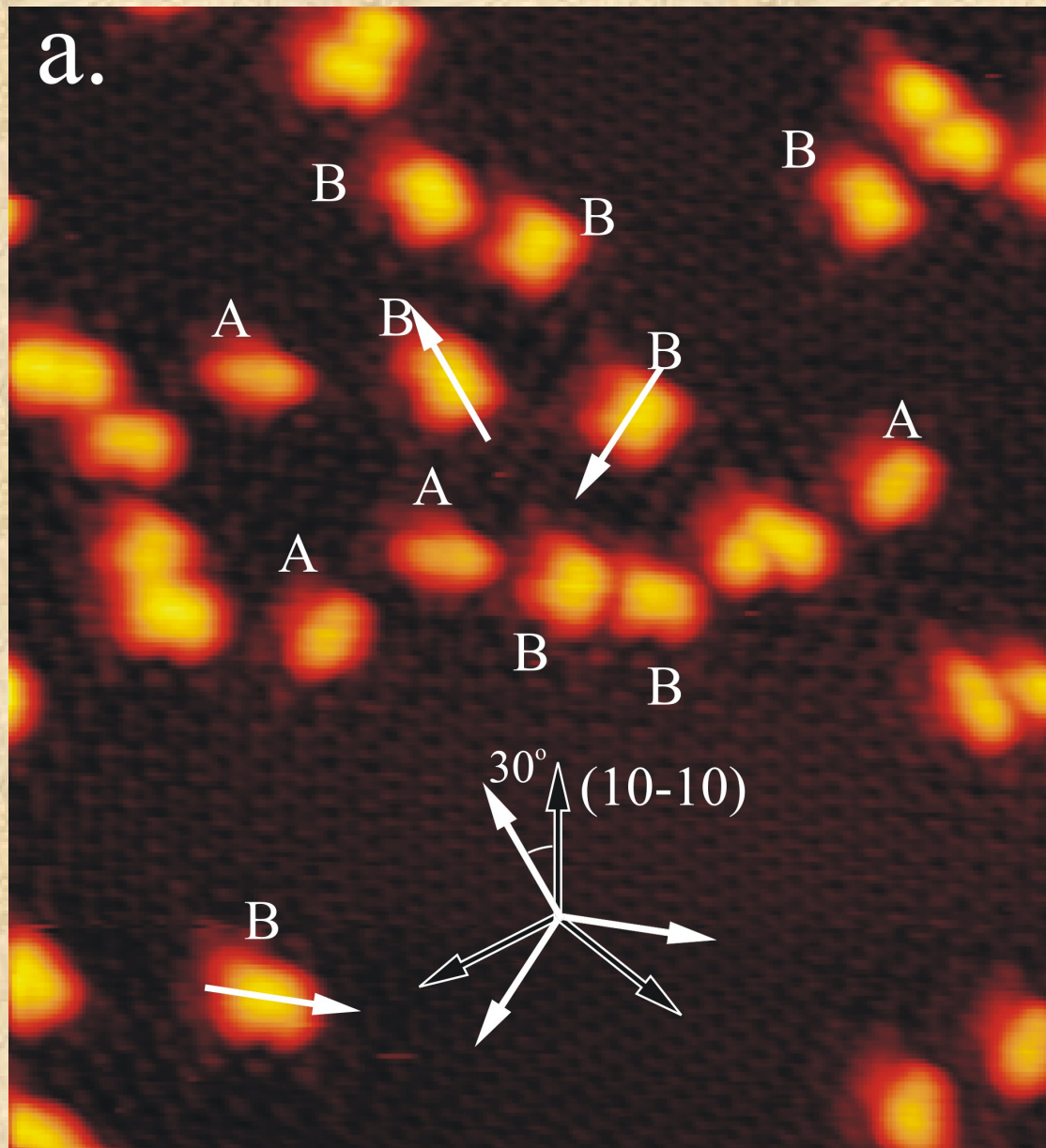


Dimer formation

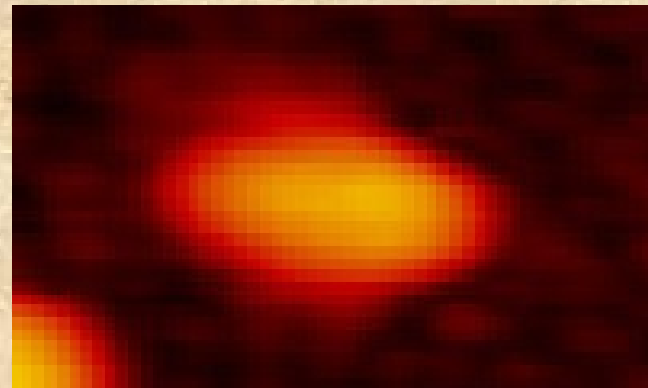


H-Dimers on graphite

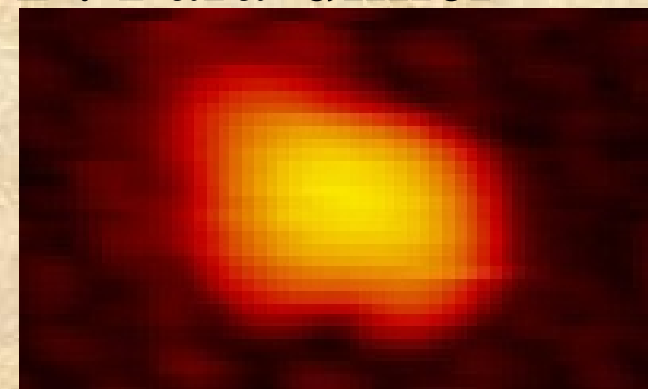
103 x 114 Å²



A: Ortho-dimer



B: Para-dimer

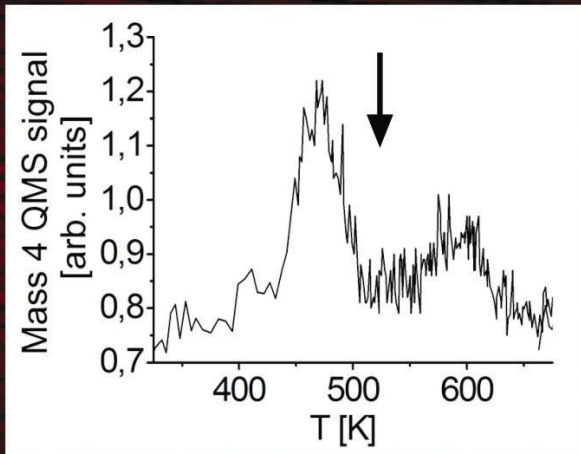


$V_t = 884 \text{ mV}$, $I_t = 0.16 \text{ nA}$

Dimers after Anneal

103 x 114 Å²

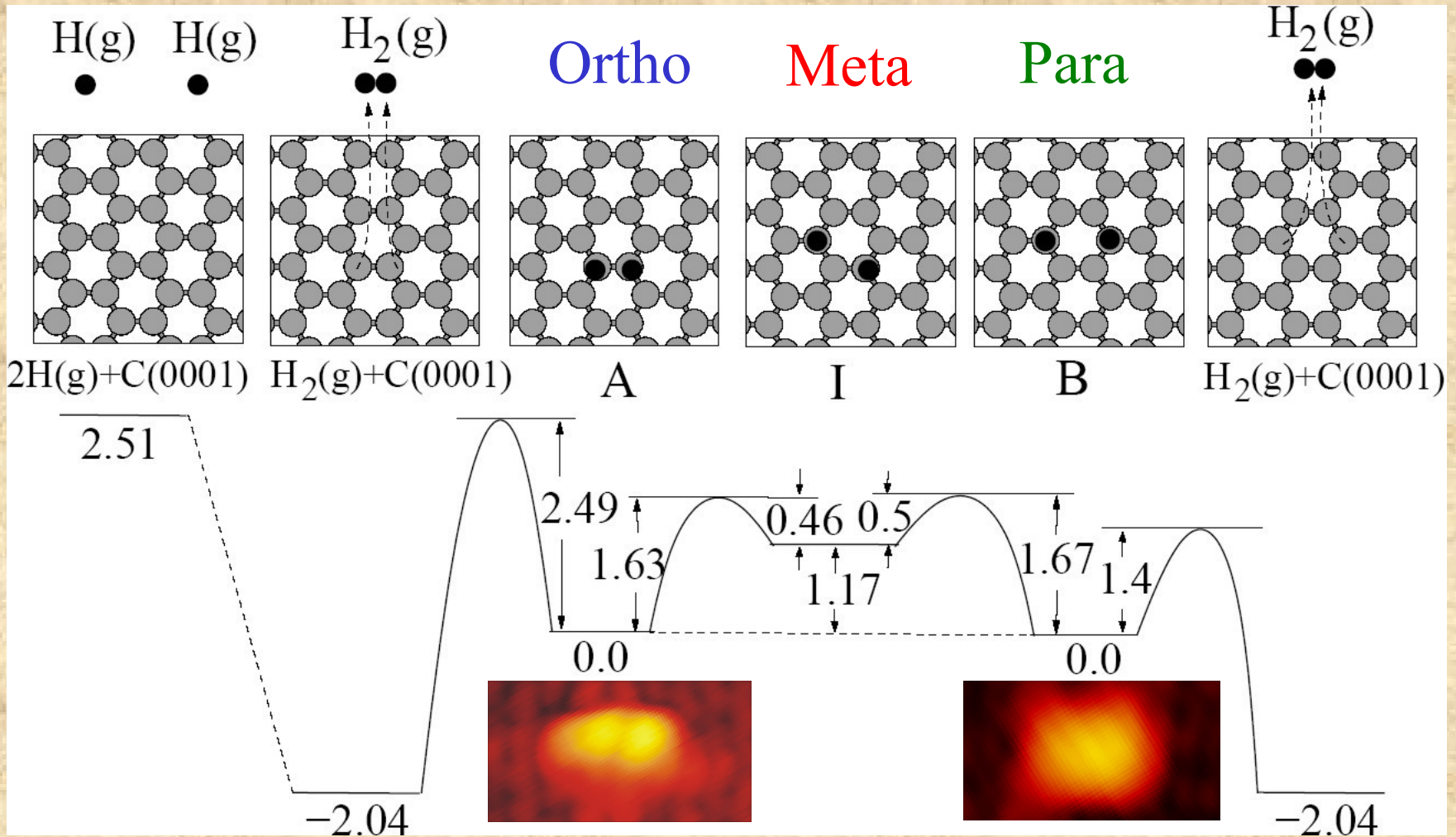
80 x 72 Å²



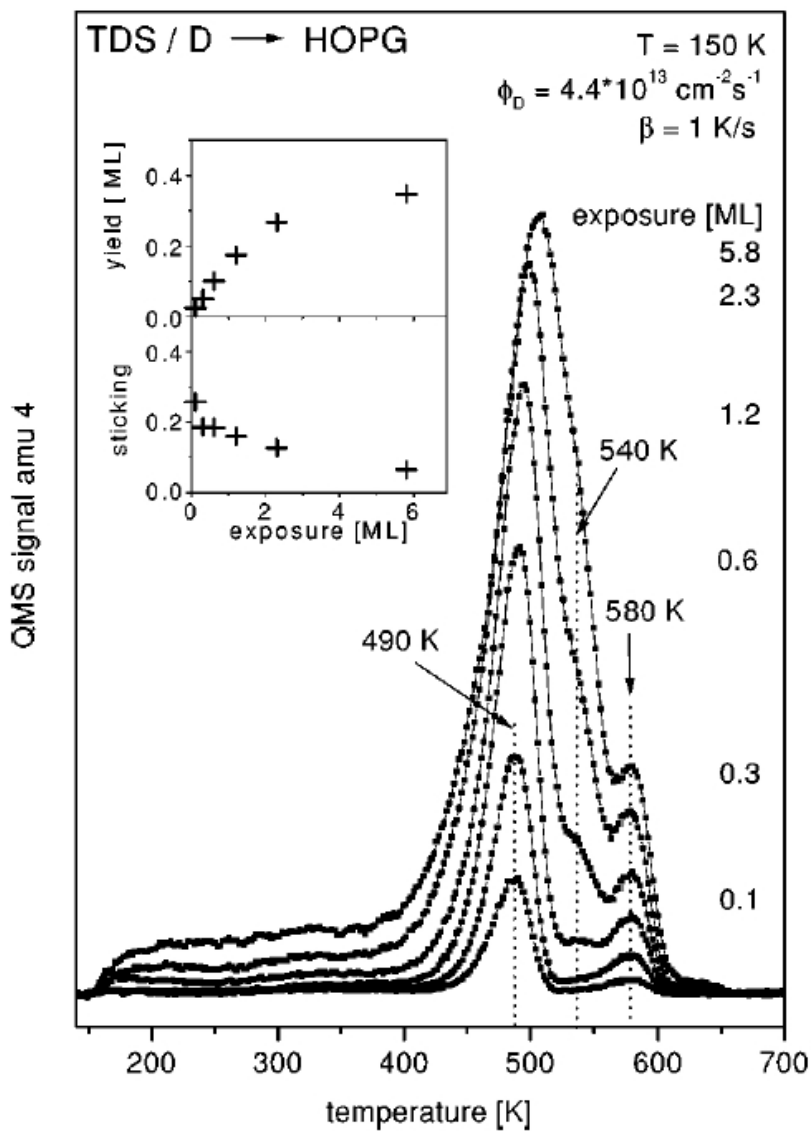
$V_t = 884 \text{ mV}$, $I_t = 0.19 \text{ nA}$

$V_t = 884 \text{ mV}$, $I_t = 0.36 \text{ nA}$

Recombination pathways



Explaining the TPD?



$$\frac{d\Theta}{dt} = -k_0 e^{-E_B/k_B T} \Theta^n$$

$n=1 \Rightarrow$ First order desorption

Barrier to diffusion: 1.3 eV

Barrier to desorption: 0.9 eV

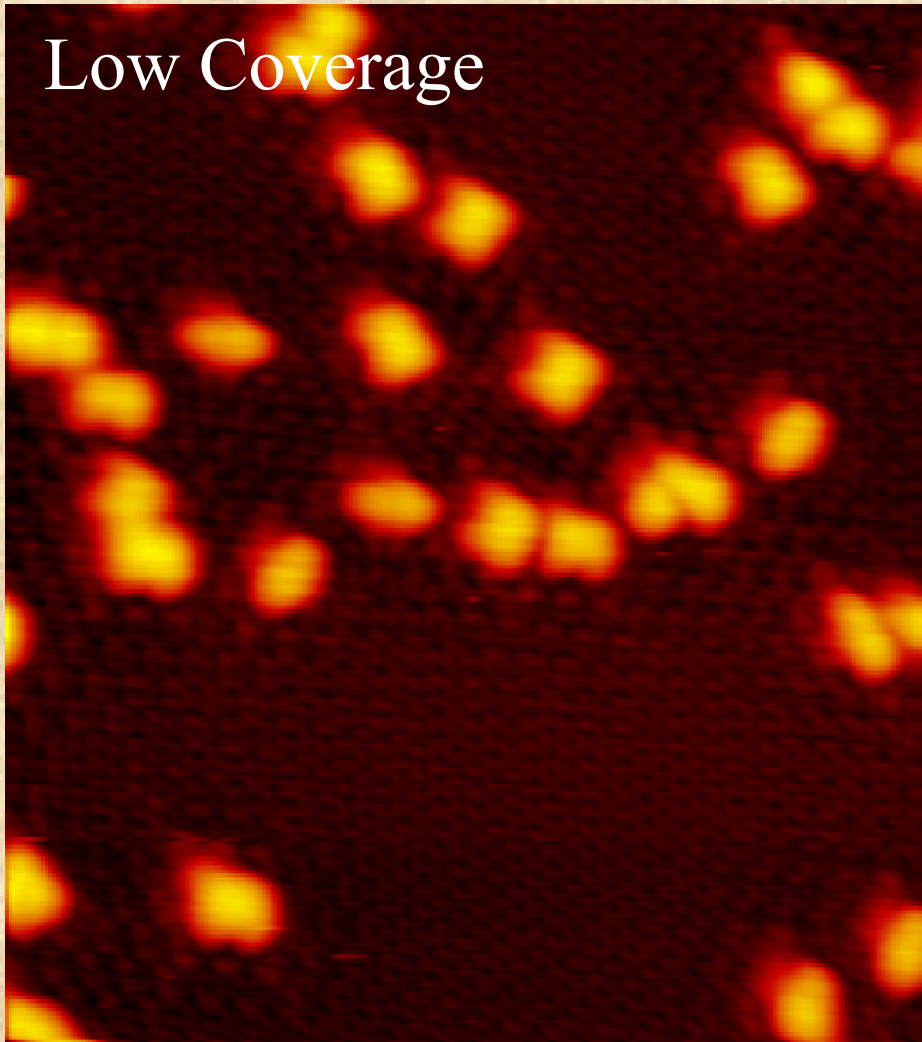
490 K \Rightarrow 1.4 eV

580 K \Rightarrow 1.6 eV

H on HOPG

171 x 155 Å²

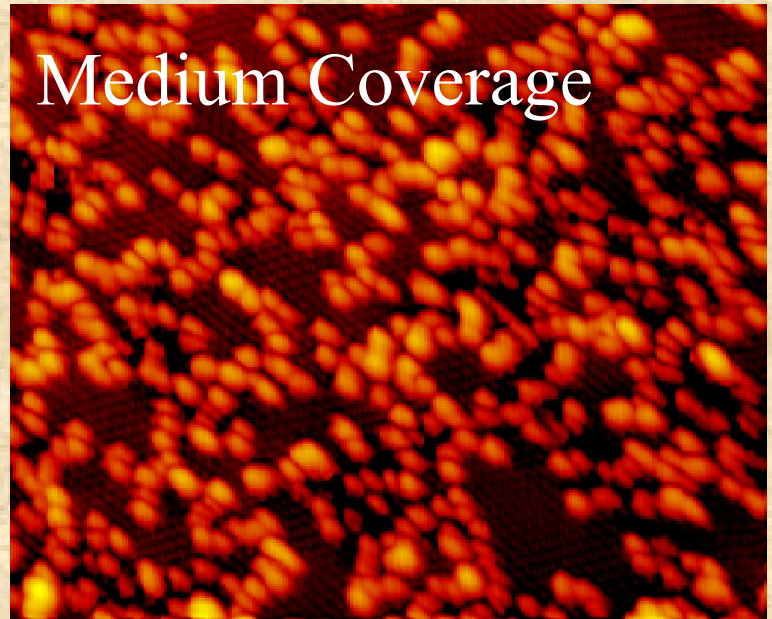
Low Coverage



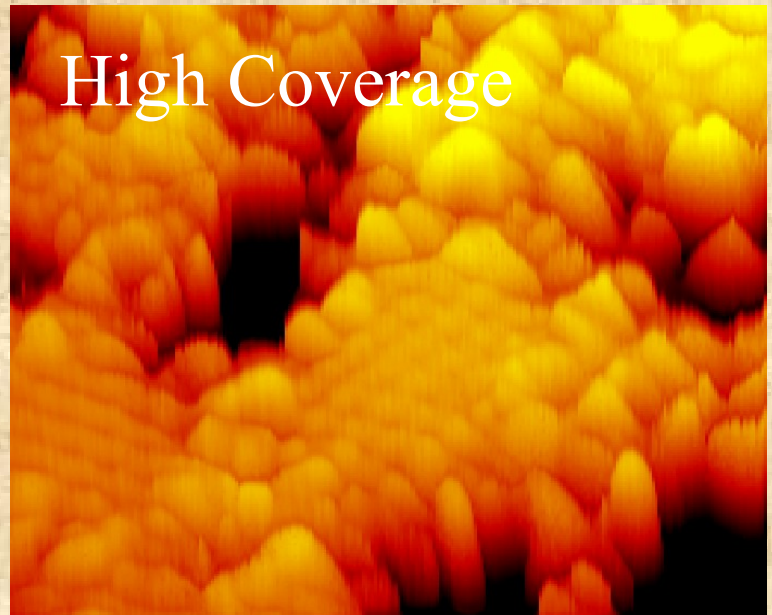
103 x 114 Å²

$V_t \sim 800\text{mV}$, $I_t \sim 0.15\text{-}0.2\text{nA}$

Medium Coverage

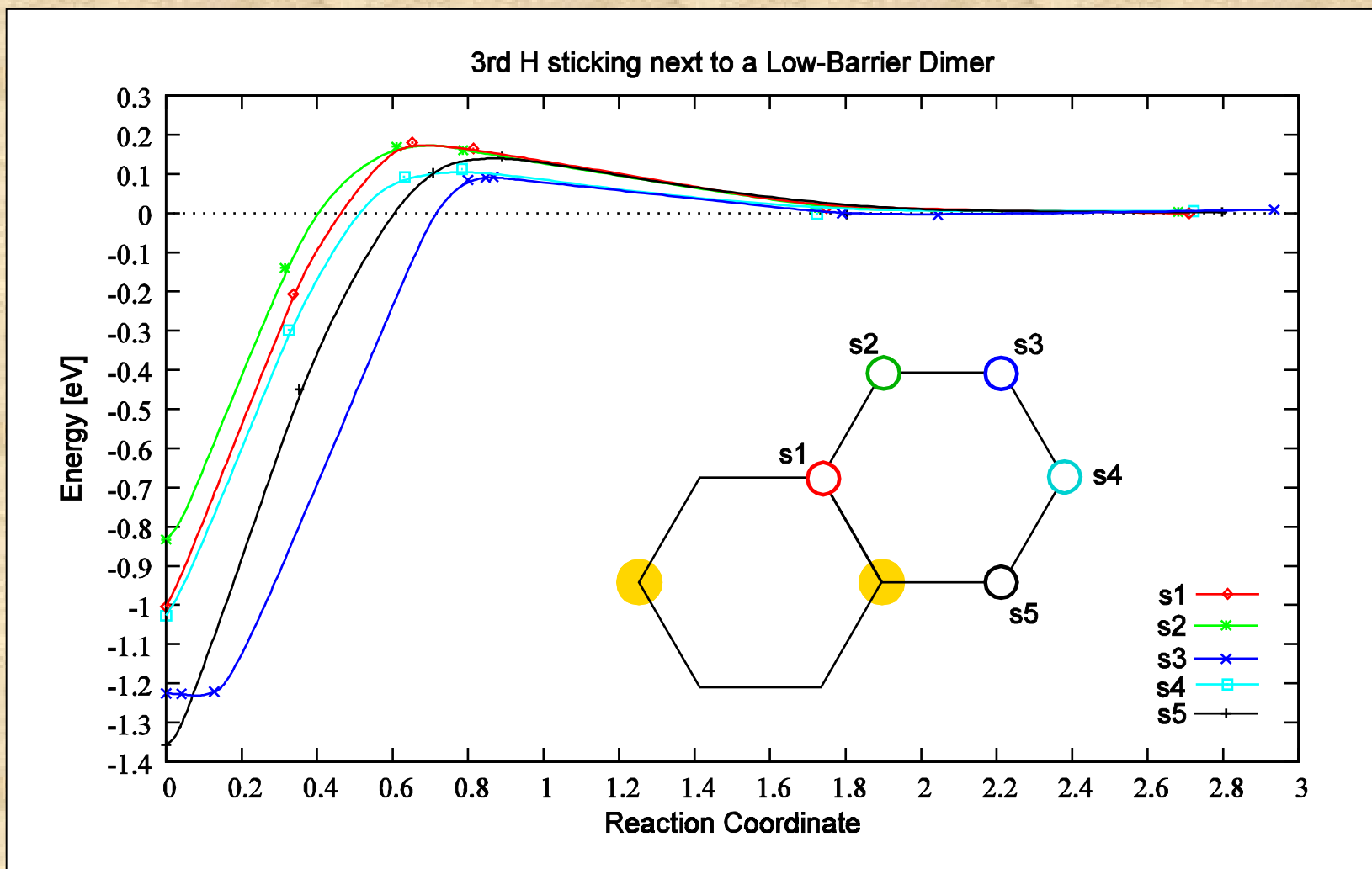


High Coverage

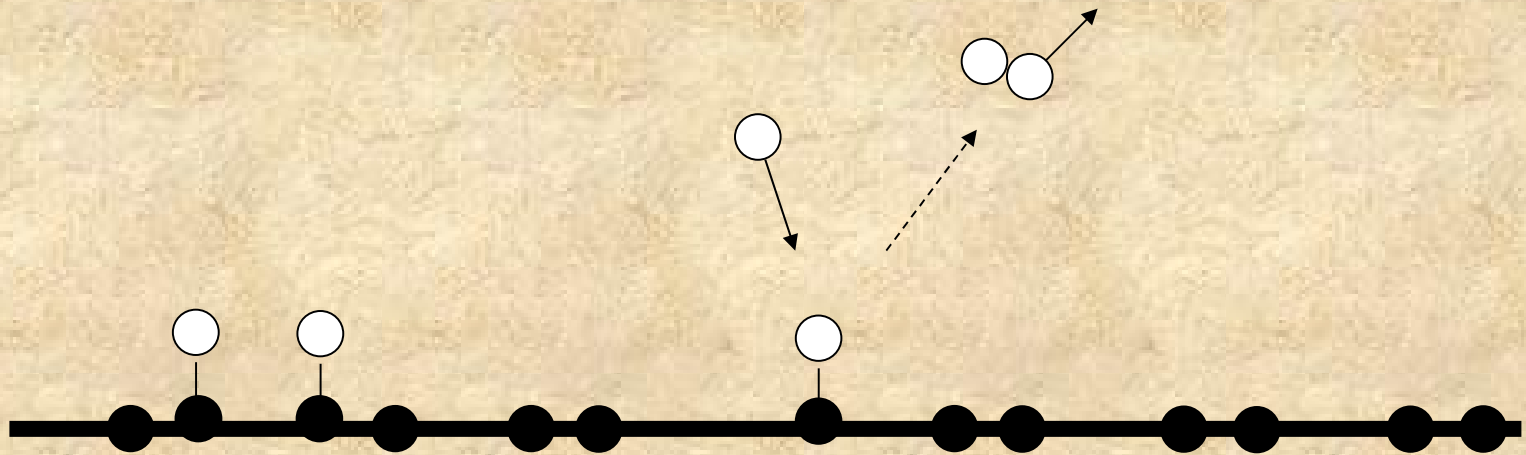


80 x 72 Å²

Preferential sticking and clustering



Eley Rideal - Abstraction



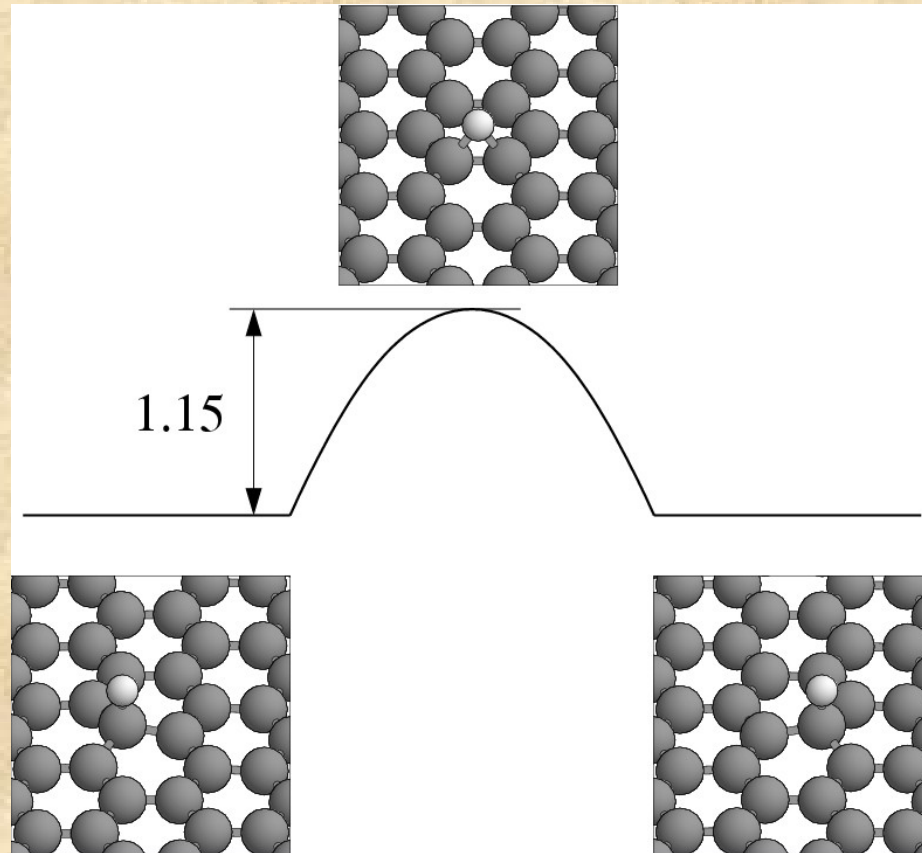
Jeloaica & Sidis (2001)

Sha et al (2002)

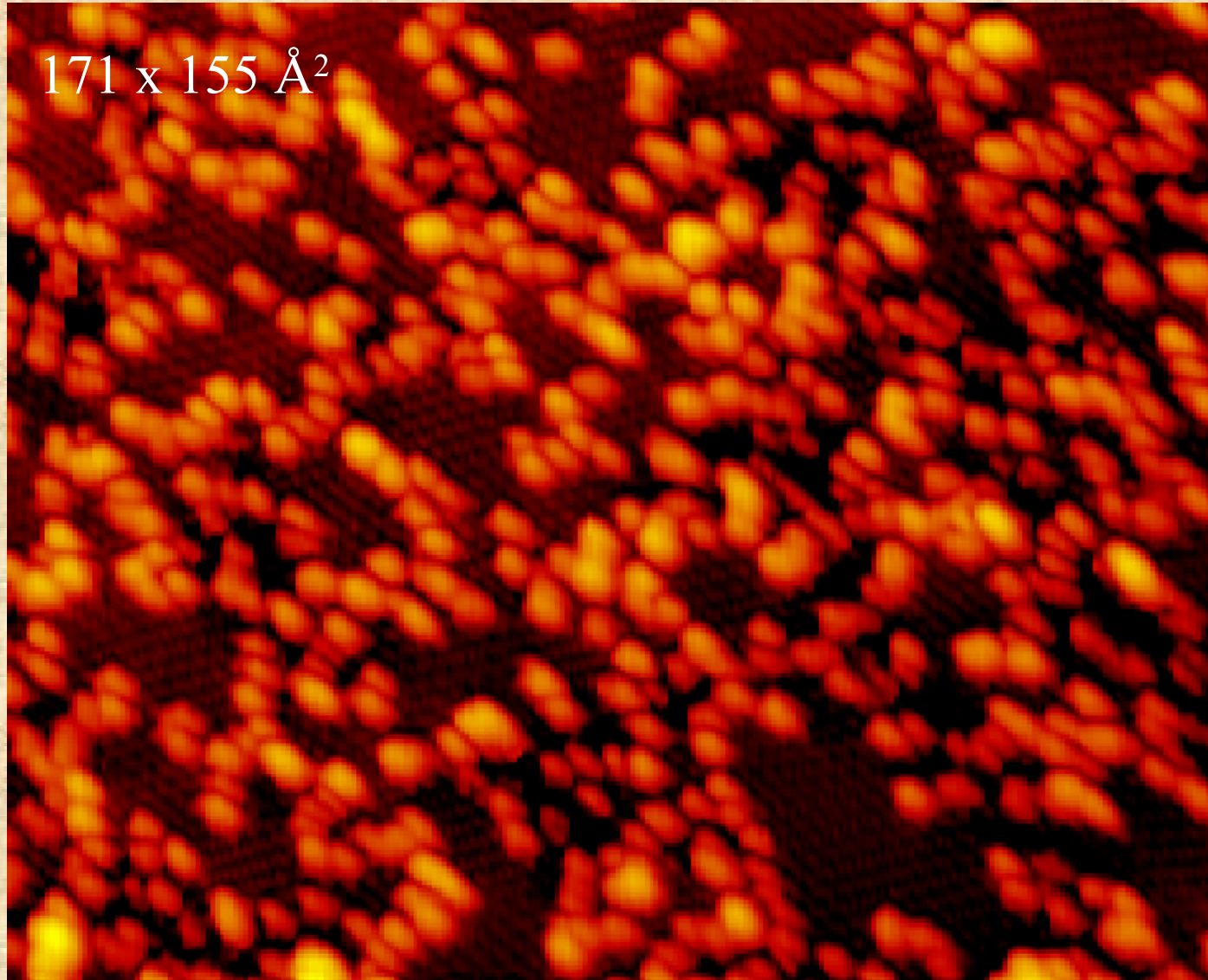
Zecho et al (2002)

Diffusion

Barrier to diffusion for an isolated H atom: 1.15 eV



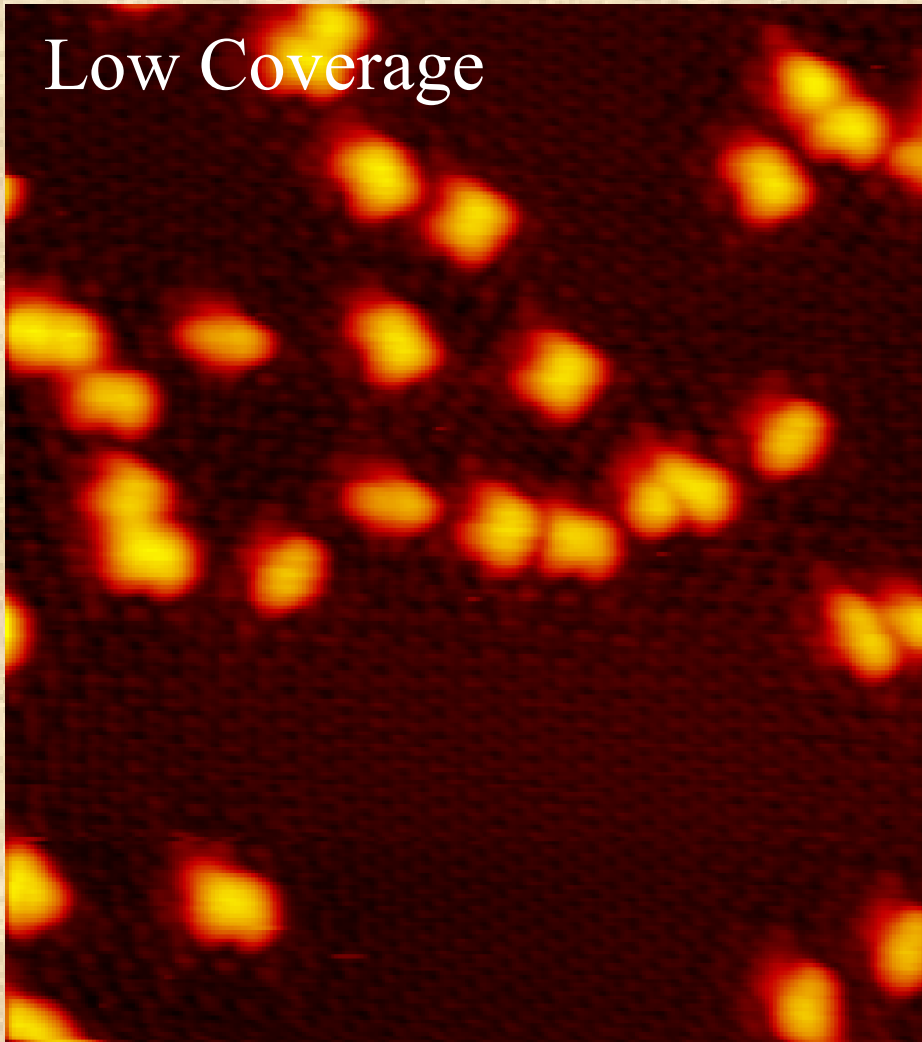
Random adsorption



H on HOPG

171 x 155 Å²

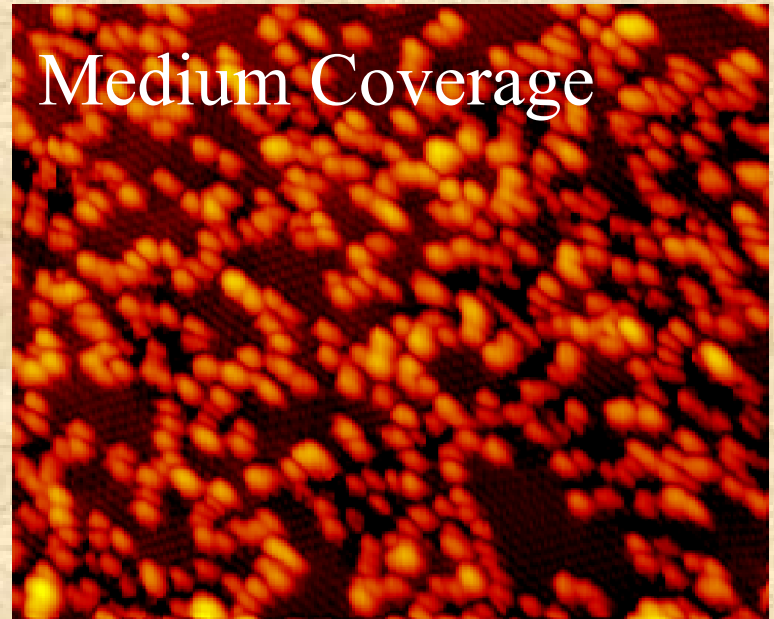
Low Coverage



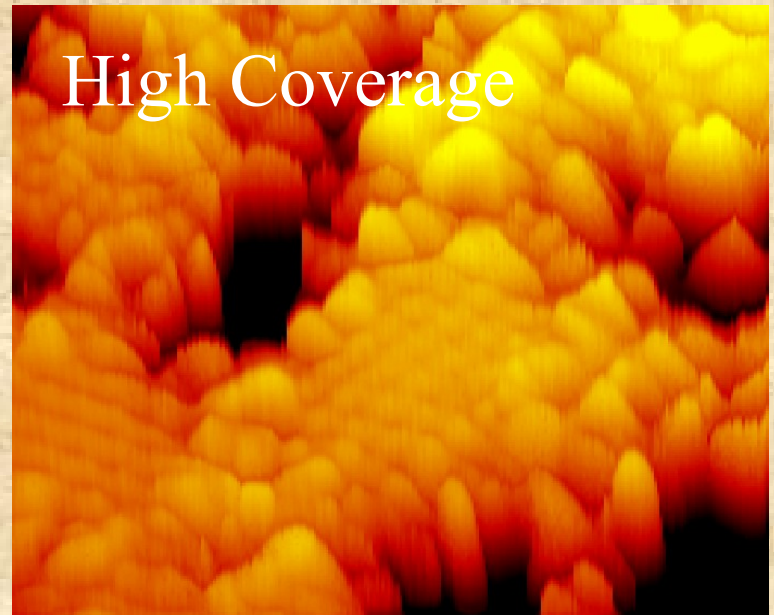
103 x 114 Å²

$V_t \sim 800\text{mV}$, $I_t \sim 0.15\text{-}0.2\text{nA}$

Medium Coverage



High Coverage



80 x 72 Å²

Comparison to H-Si (100)

H-Si (100):

1st order desorption

Non-langmuirian adsorption

Complete pairing observed in STM down to $\Theta \sim 20\%$

Suggested mechanism:

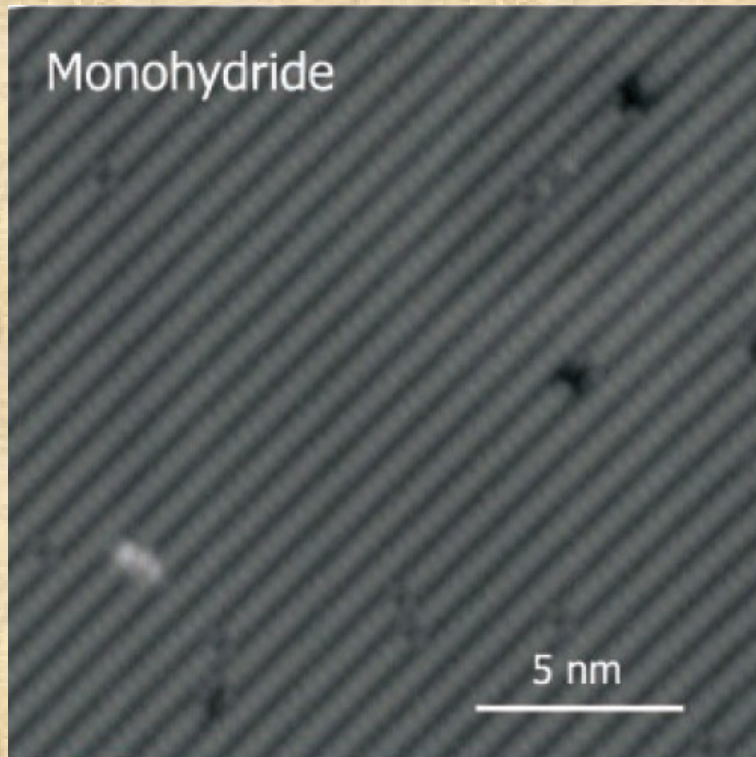
Hot atom precursor mediated adsorption

Theory predicts precursor states with $E_b \sim 1-2$ eV

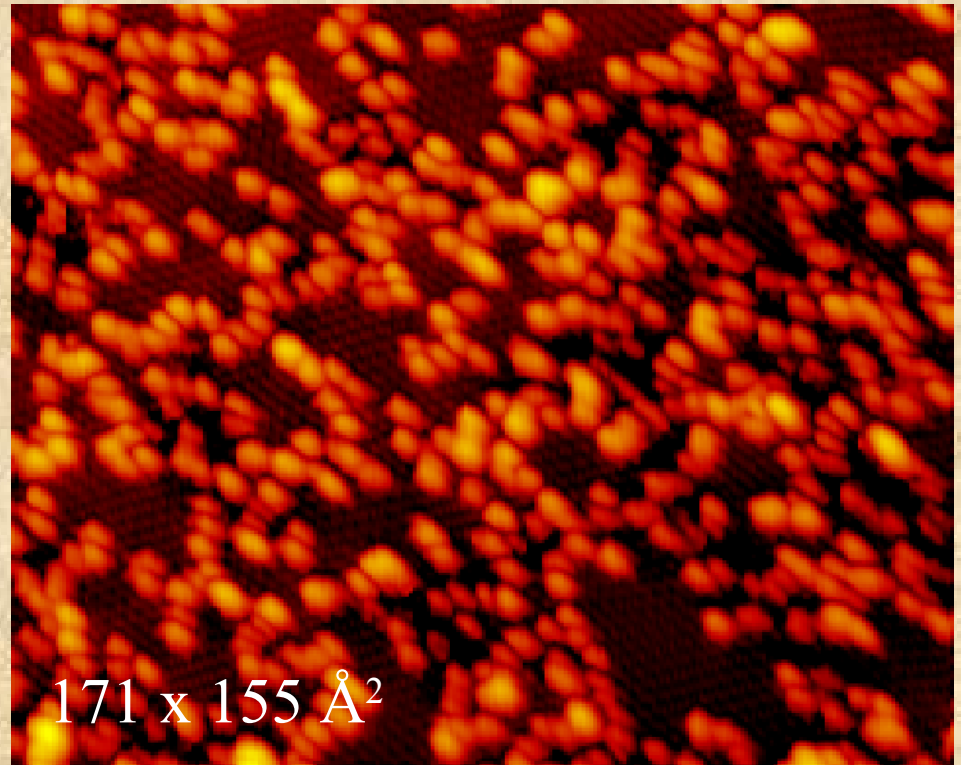
Widdra et al., PRL 74, 2074 (1995)

Tok et al, JCP 118, 3294 (2003)

Comparison to H-Si (100)

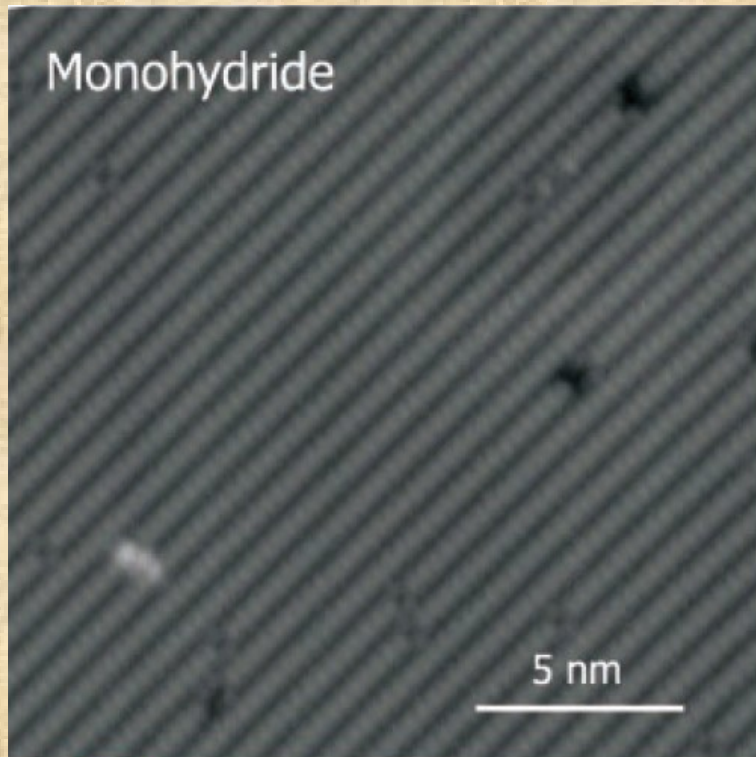


H-Si (100)

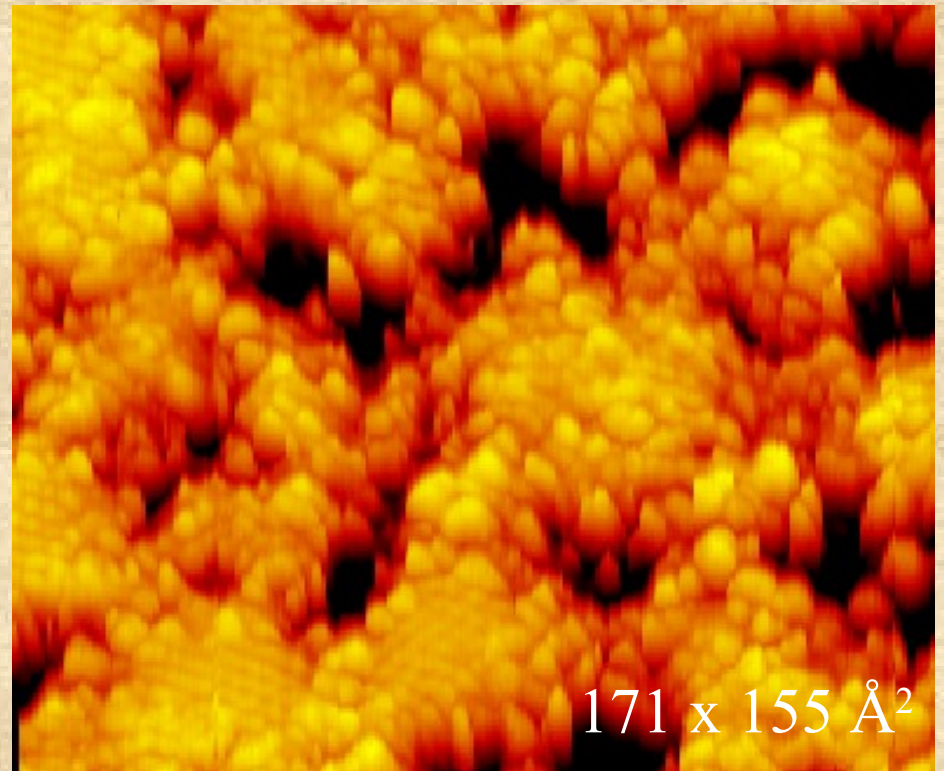


H-HOPG (0001)

Comparison to H-Si (100)

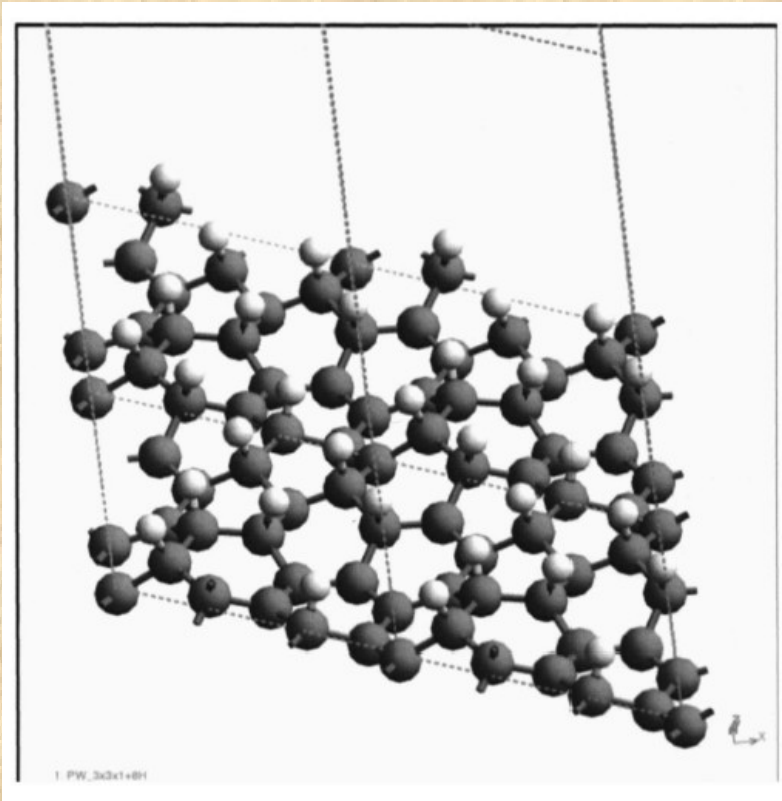


H-Si (100)

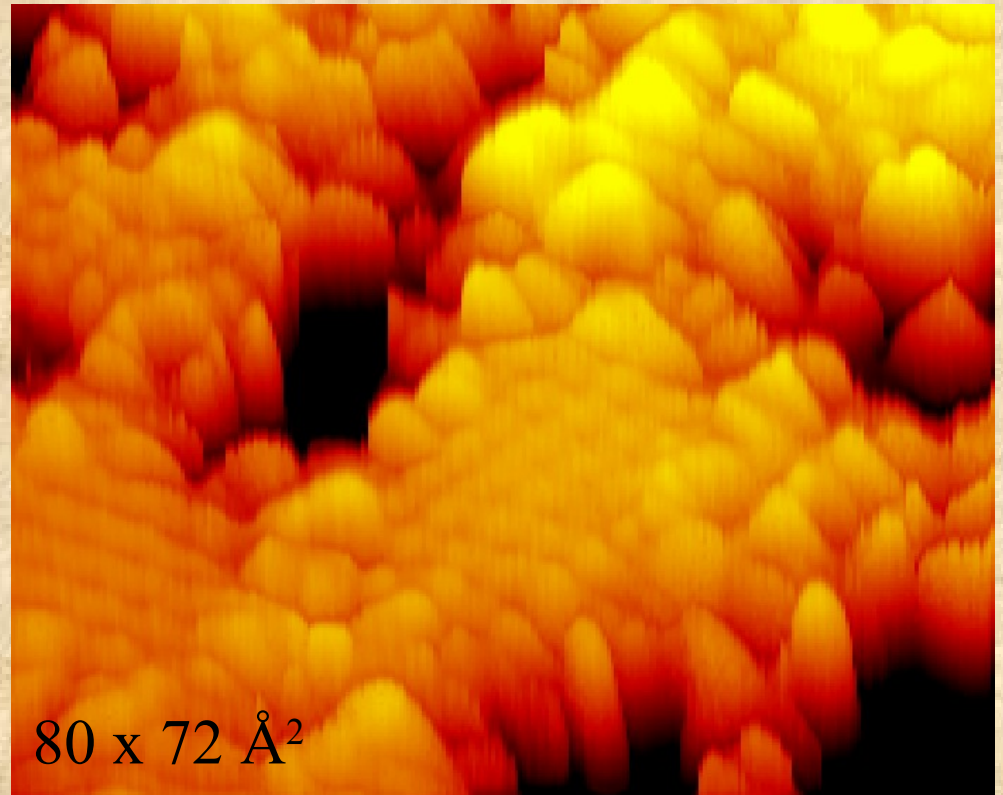


H-HOPG (0001)

High coverage



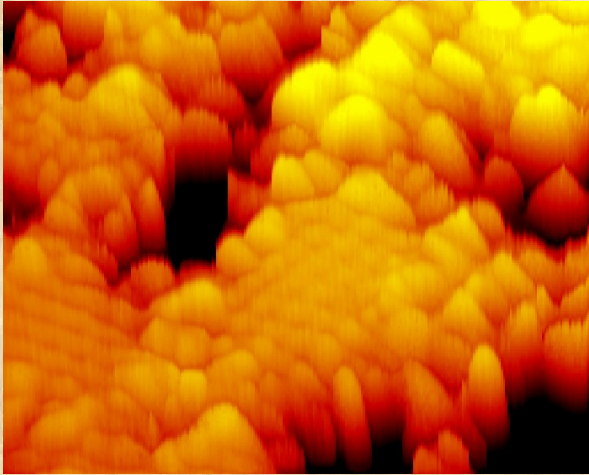
Allouche et al., J. Chem. Phys.
123, 124701 (2005)



Hornekær et al. Phys. Rev. Lett.
97, 186102 (2006)

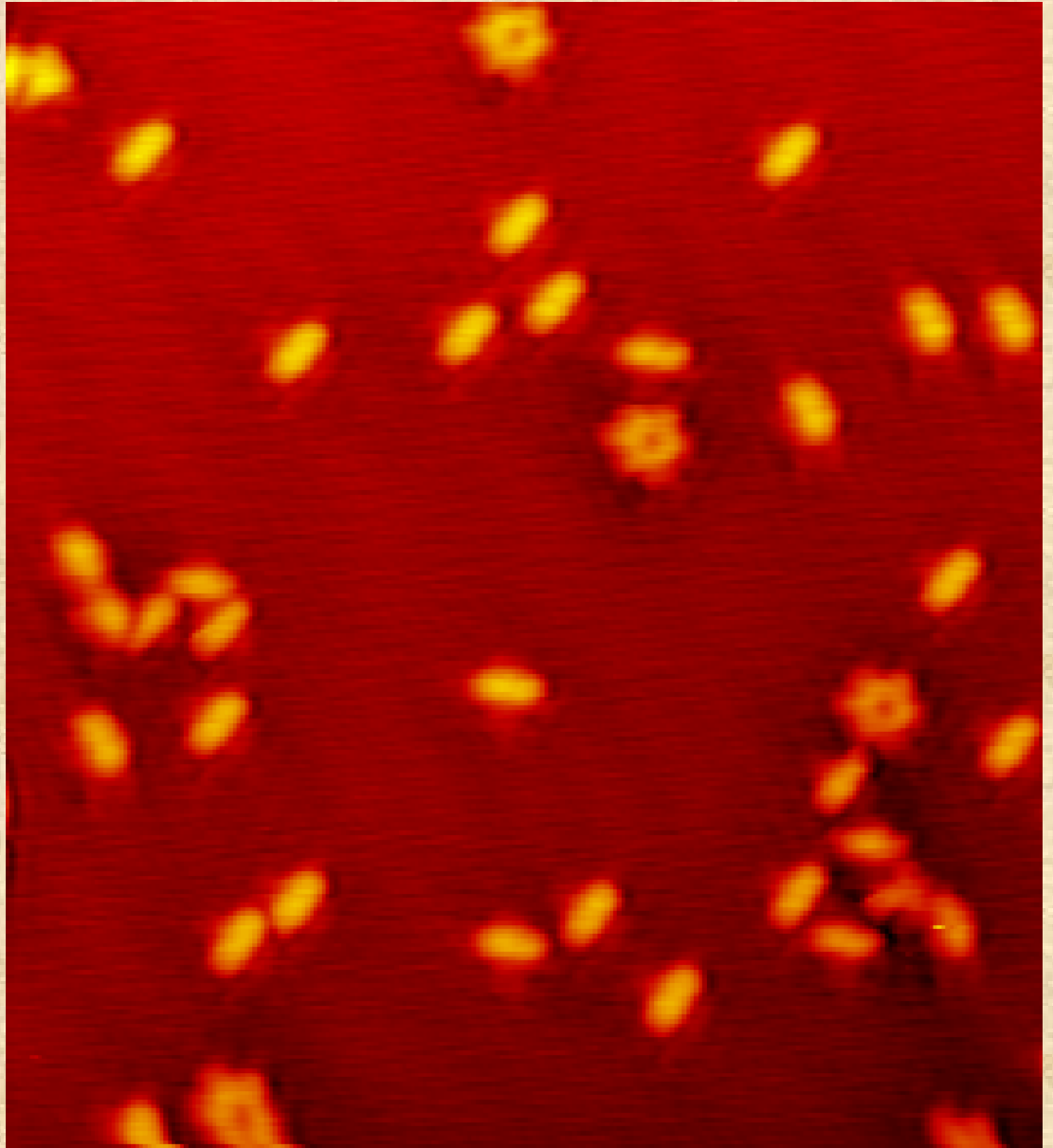
High Coverage

$V_t = -1.05$ V, $I_t = -0.55$ nA

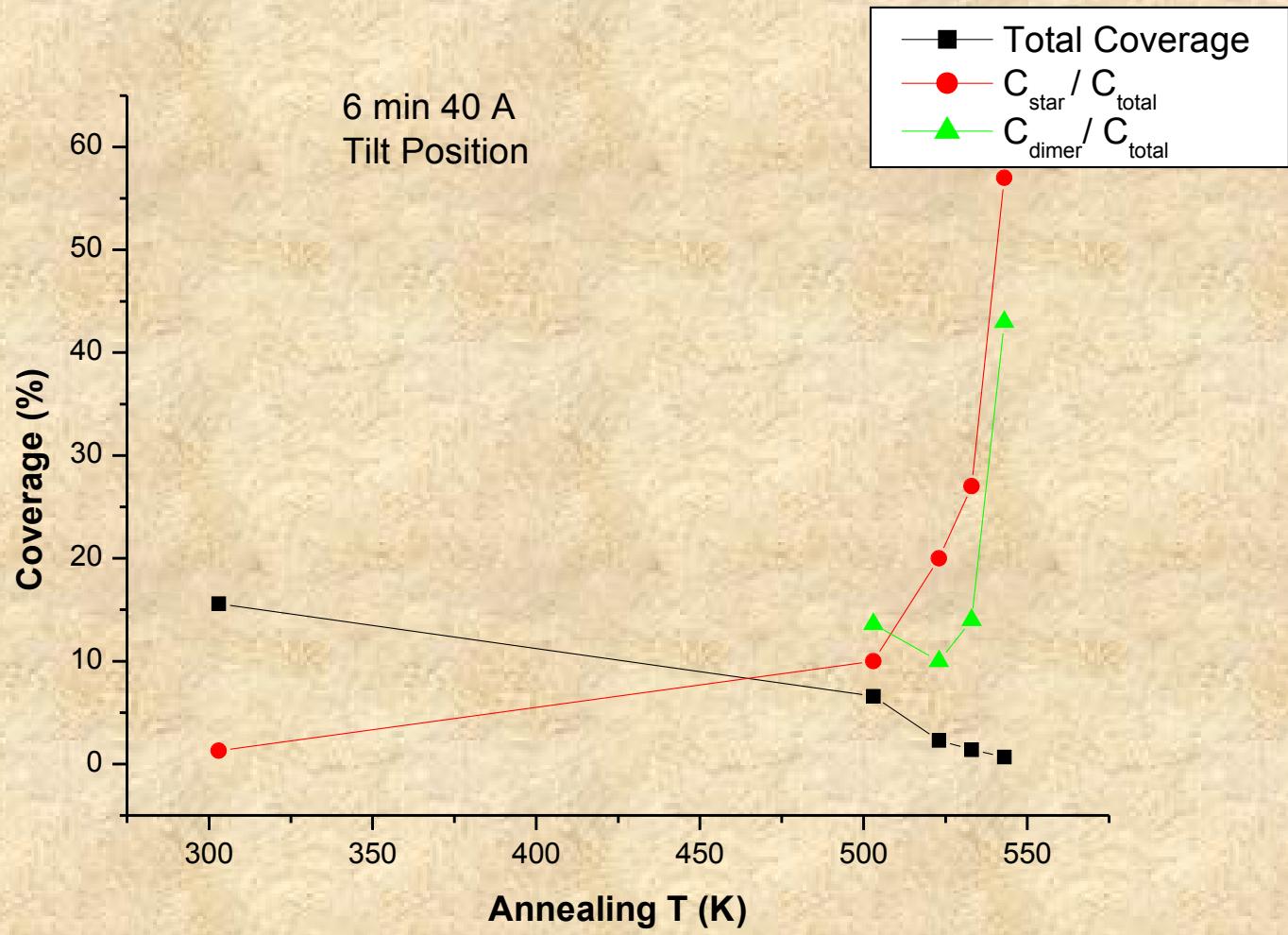


$80 \times 72 \text{ \AA}^2$

525K anneal

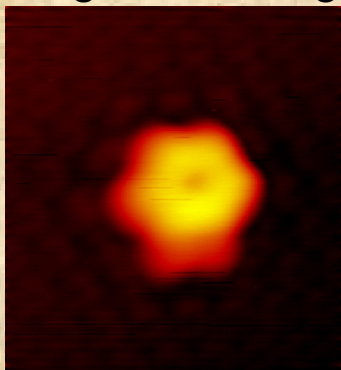


$V_t = -1.05$ V, $I_t = -0.18$ nA, $171 \times 155 \text{ \AA}^2$

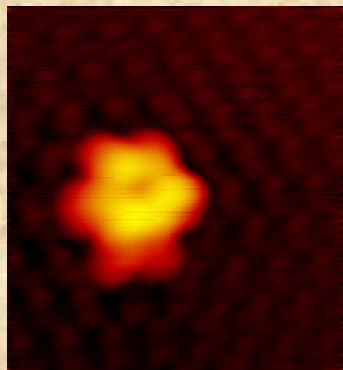


Stars / trimers

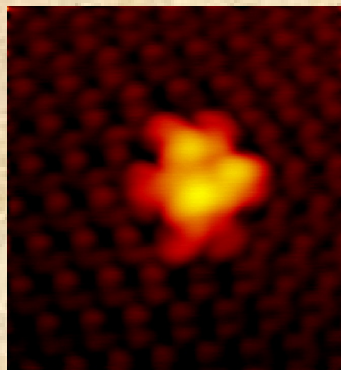
Negative voltages:



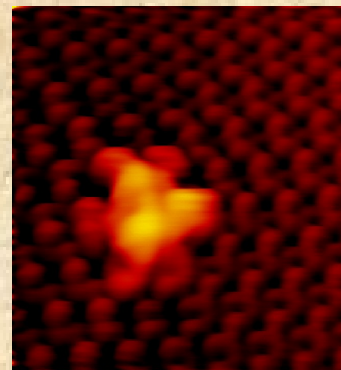
I=-0.16nA, V=-874mV
0512020424



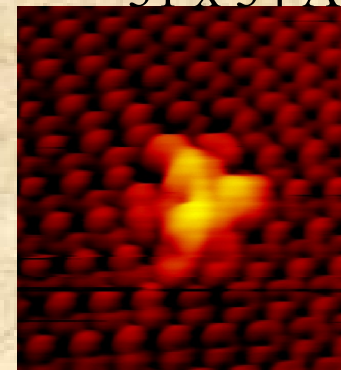
I=-0.15nA, V=-309mV
0512020422



I=-0.16nA, V=-109mV
0512020408



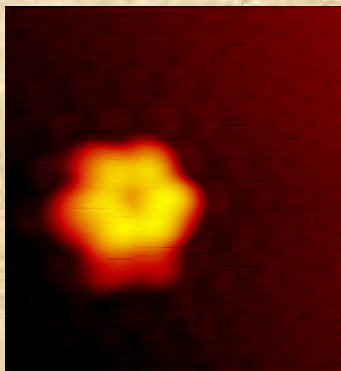
I=-0.16nA, V=-46mV
05120204210



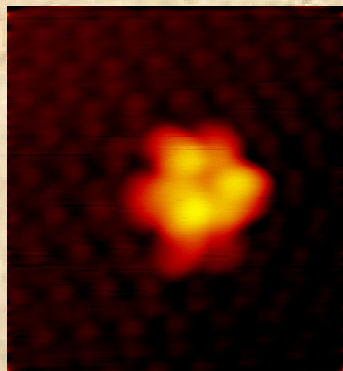
31 x 34 Å²

I=-0.15nA, V=-23mV
0512020437

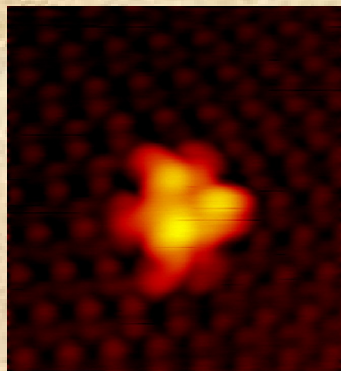
Positive voltages:



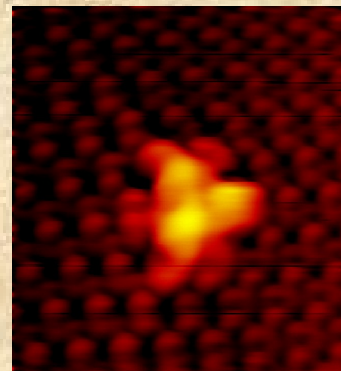
I=0.15nA, V=874mV
0512020435



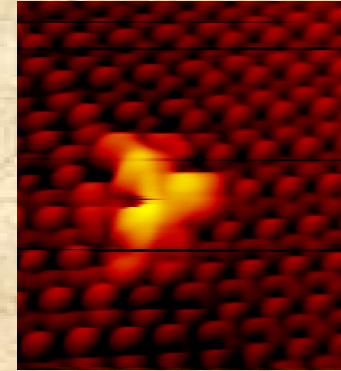
I=0.15nA, V=367mV
0512020429



I=0.15nA, V=154mV
0512020431

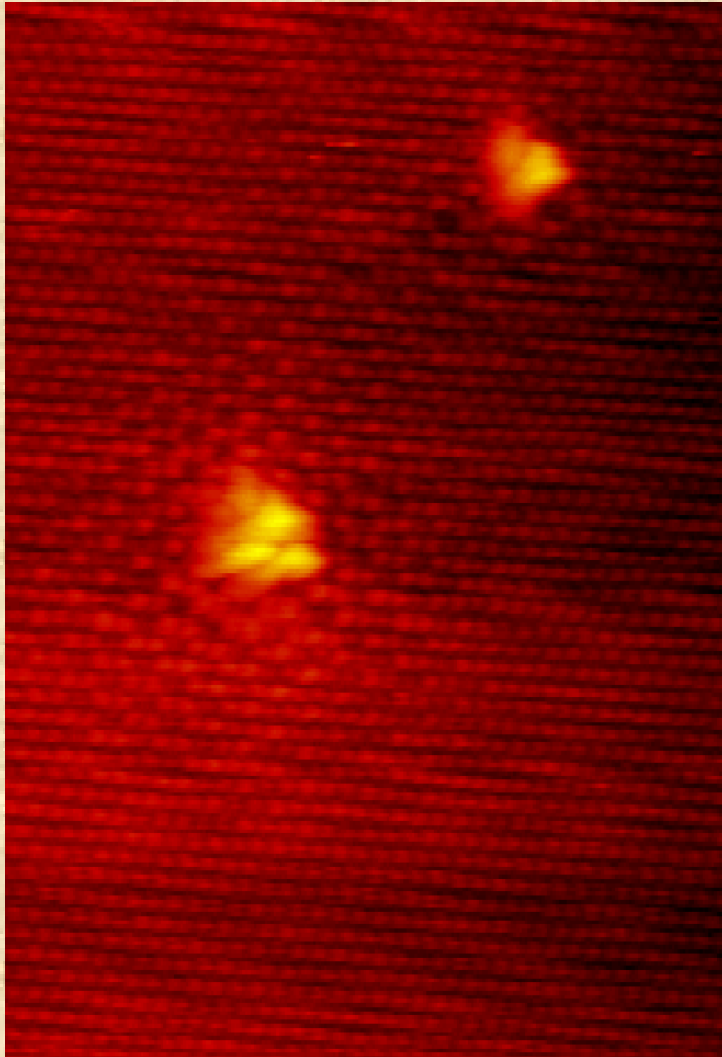


I=0.15nA, V=46mV
0512020434



I=0.15nA, V=23mV
0512020438

Trimers



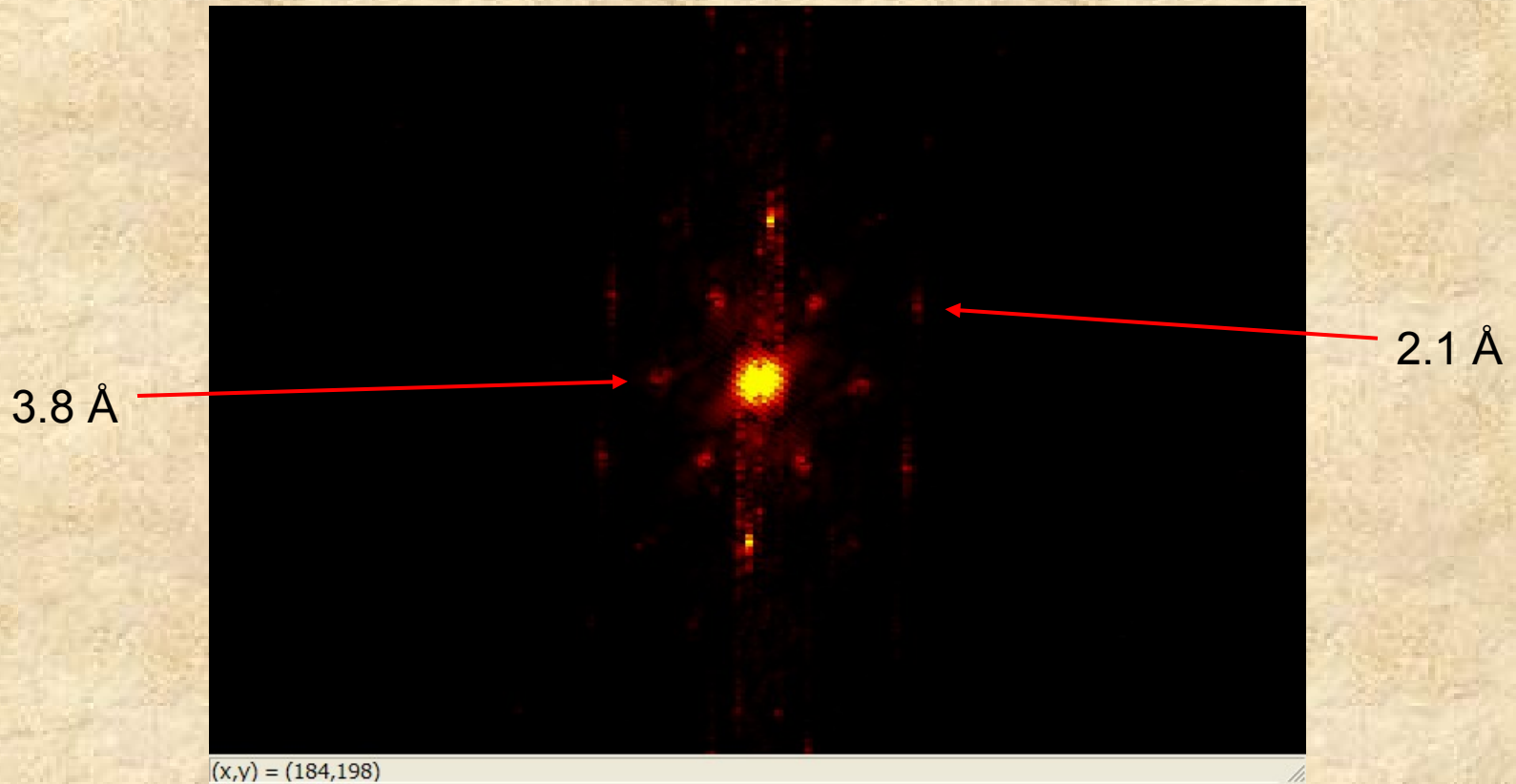
85 Å

0421-8-27

$I_t = -0.220\text{nA}$

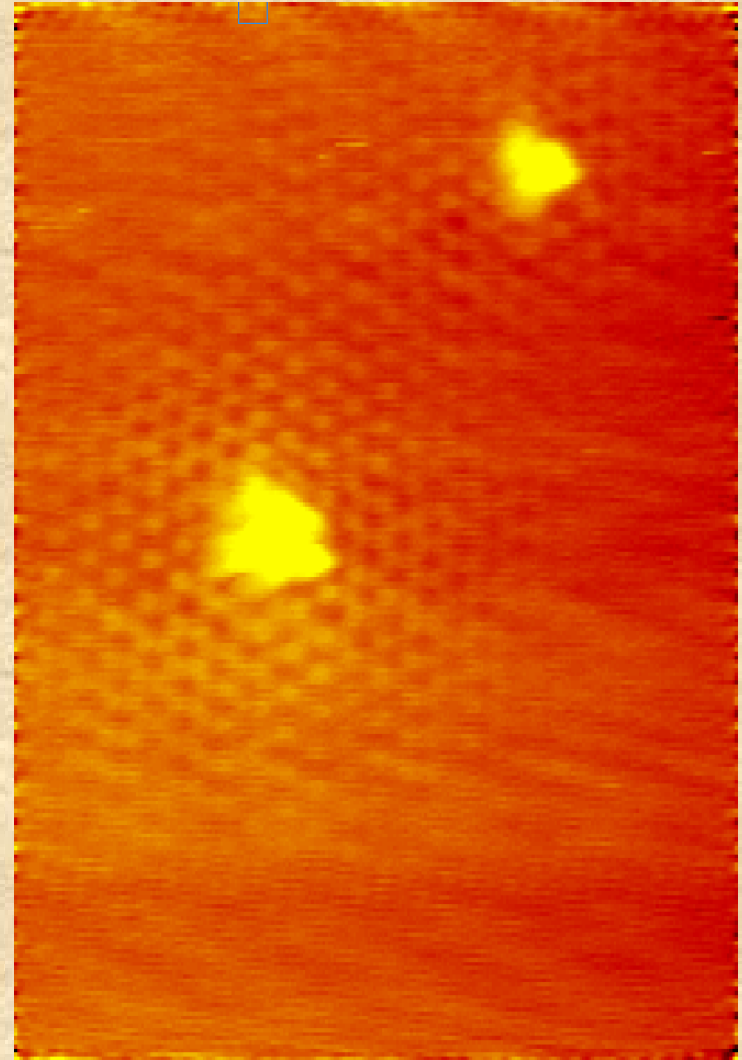
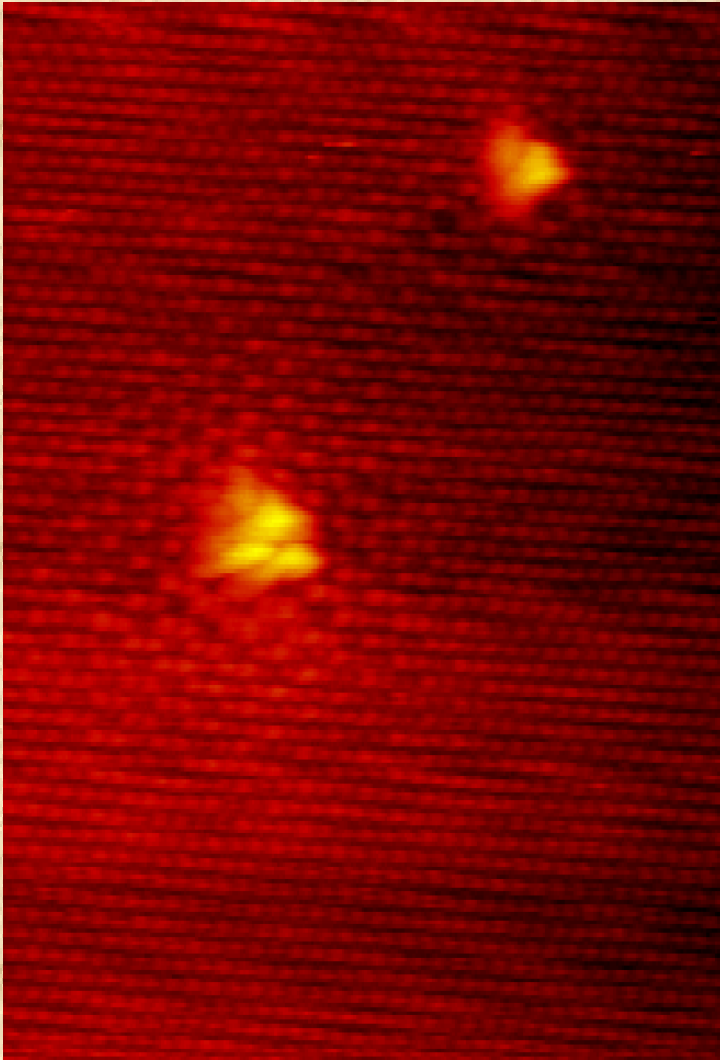
$V_t = -625\text{mV}$

Trimers - superstructure

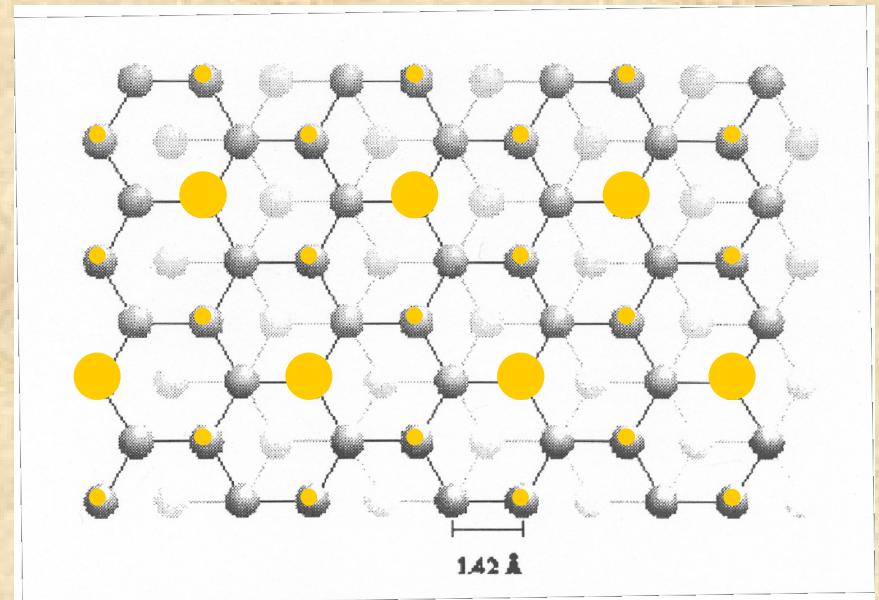
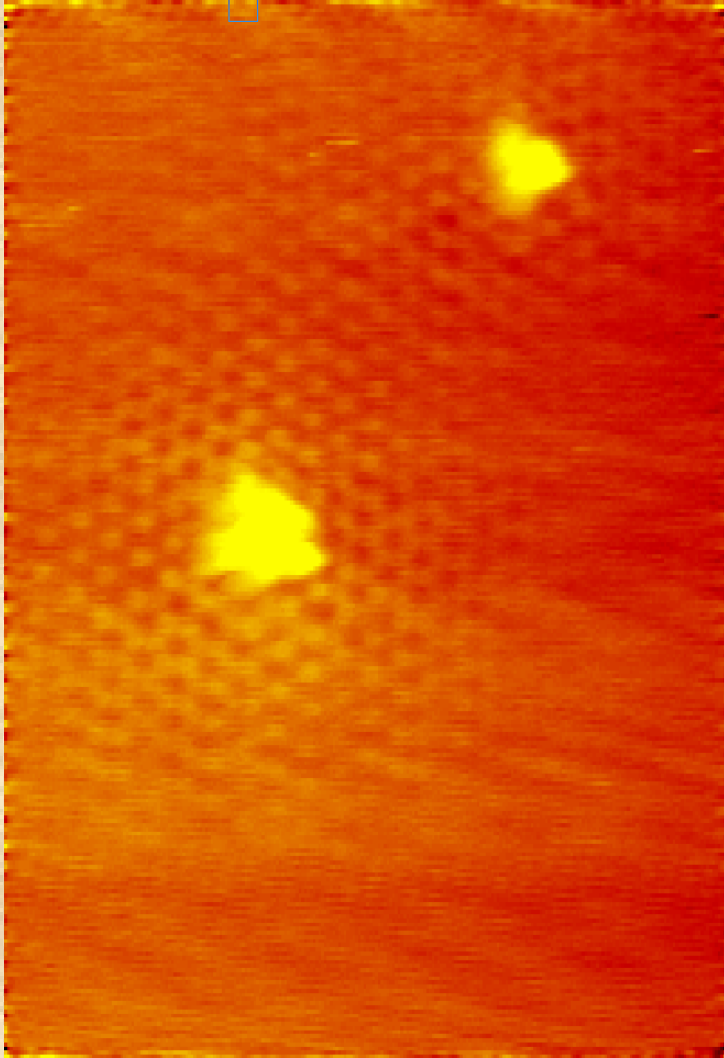


Squareroot(3) reconstruction

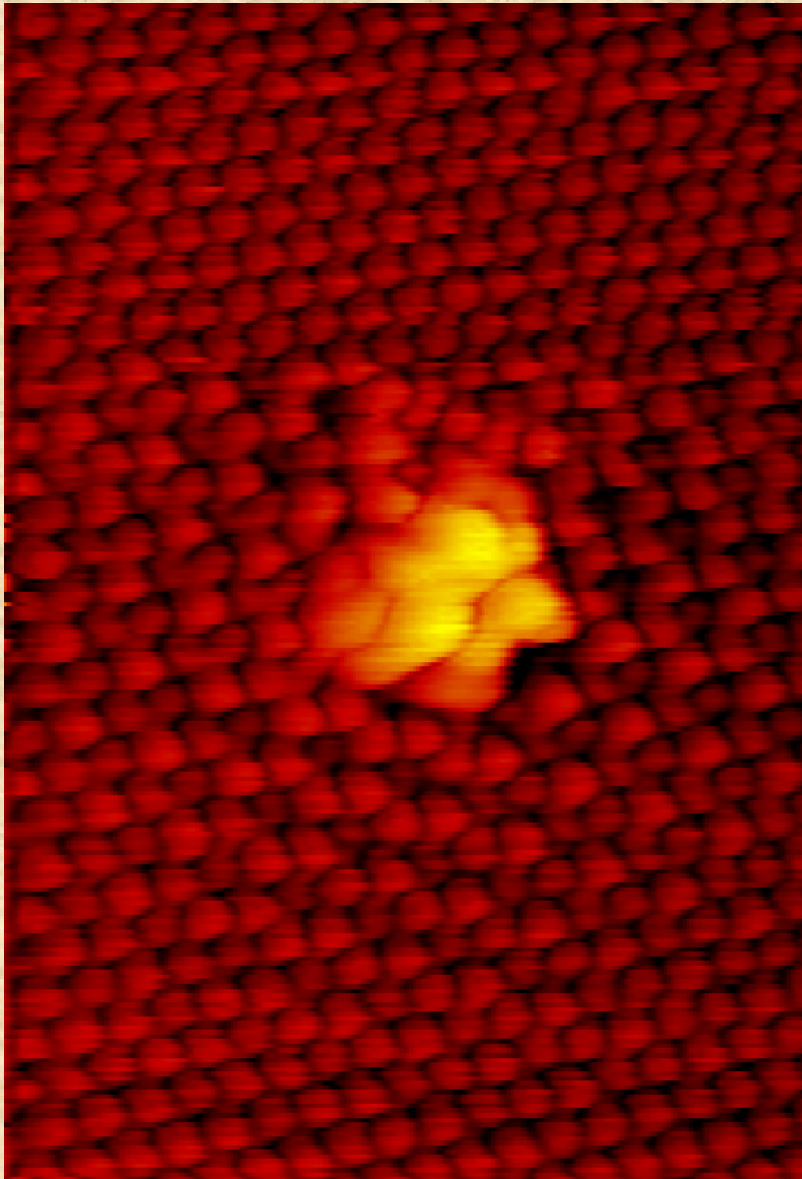
Trimers - superstructure



Trimers - superstructure



Tip induced desorption



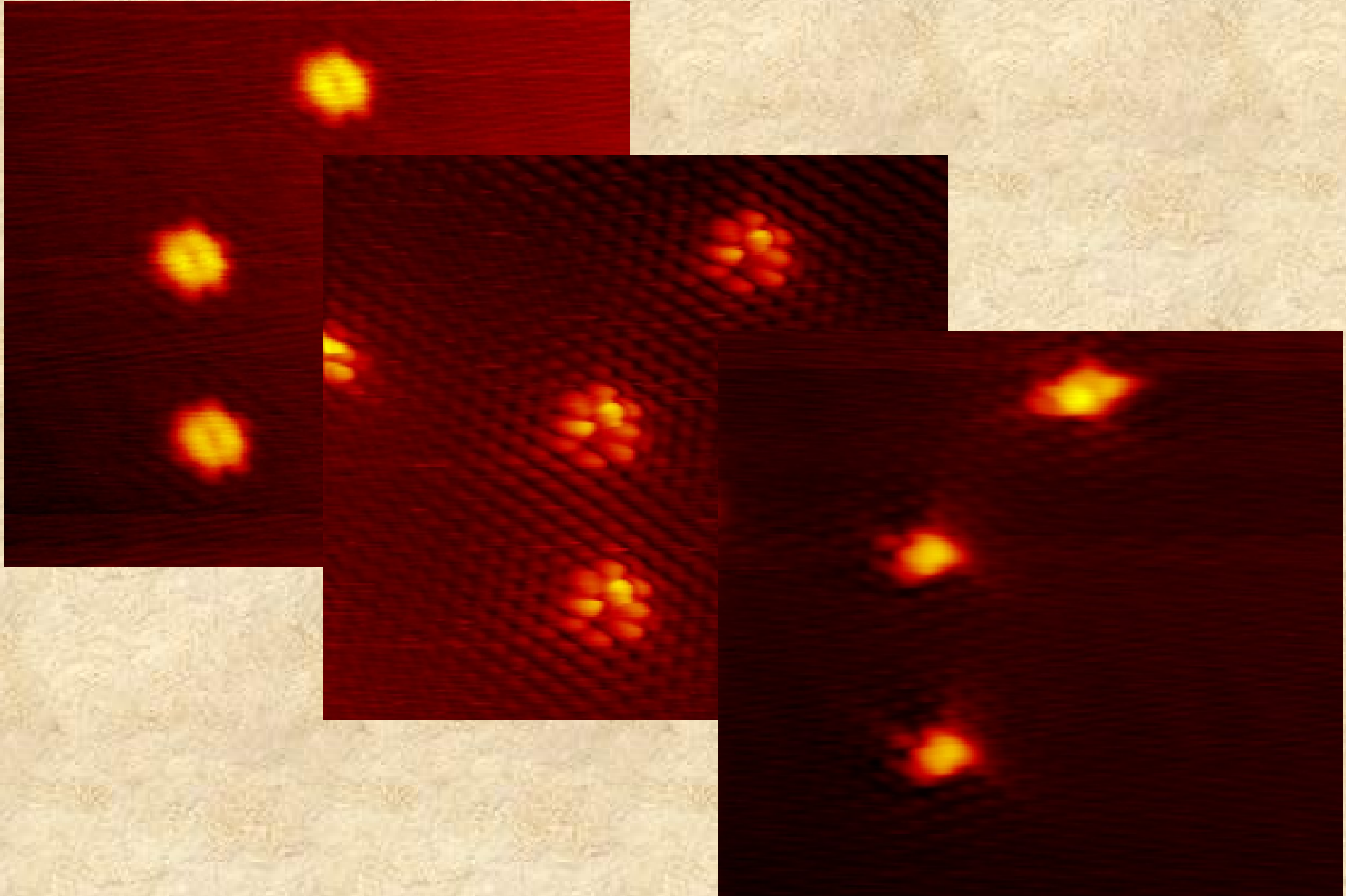
42 Å

0420-7-5

$I_t = -0.150\text{nA}$

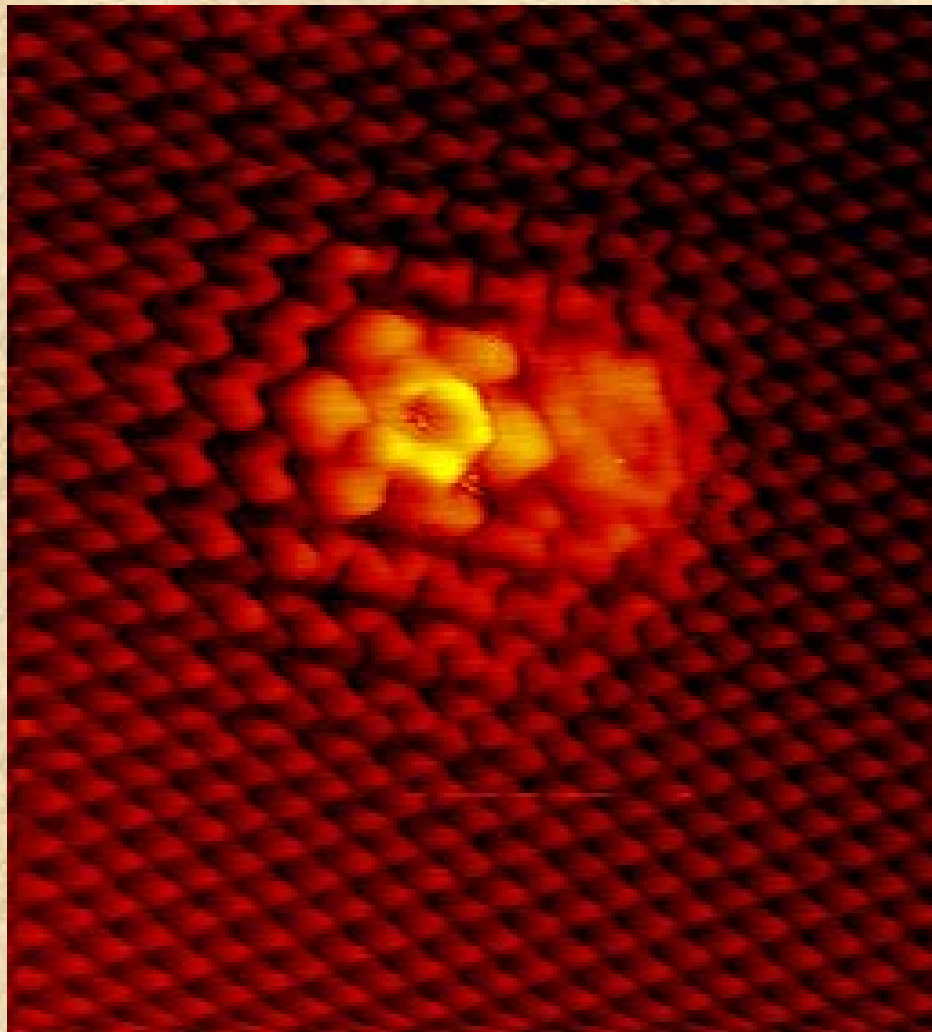
$V_t = -442\text{mV}$

Manipulation

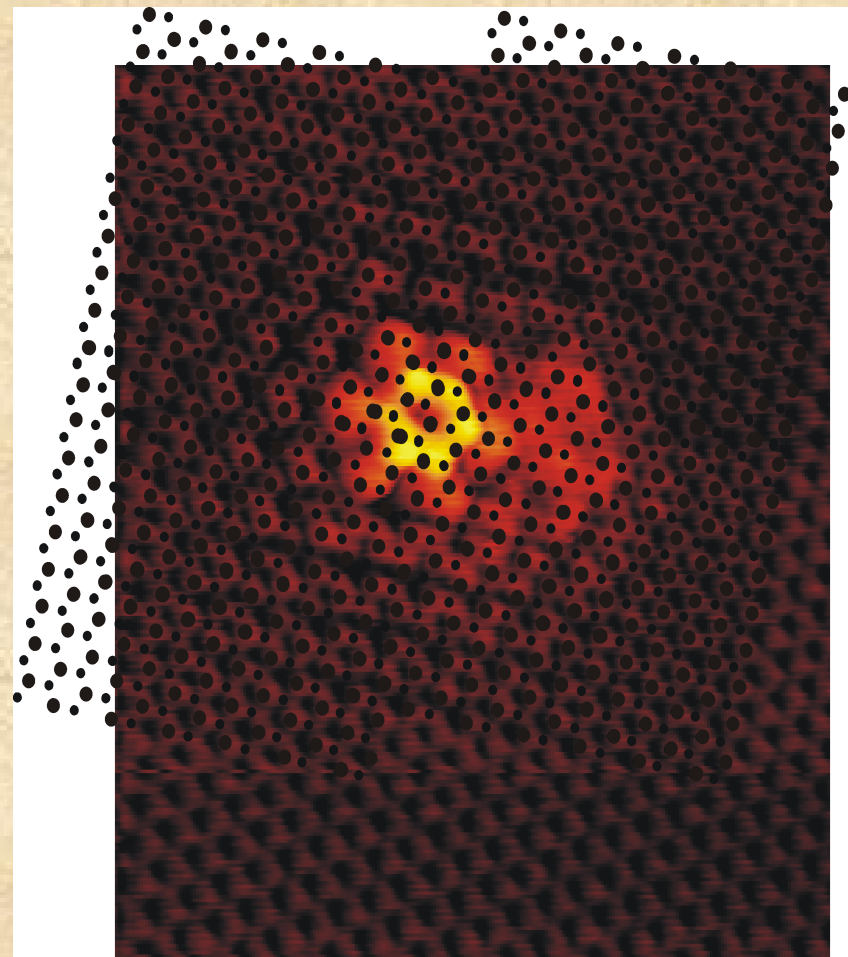


Star-structure

52 x 57 Å²

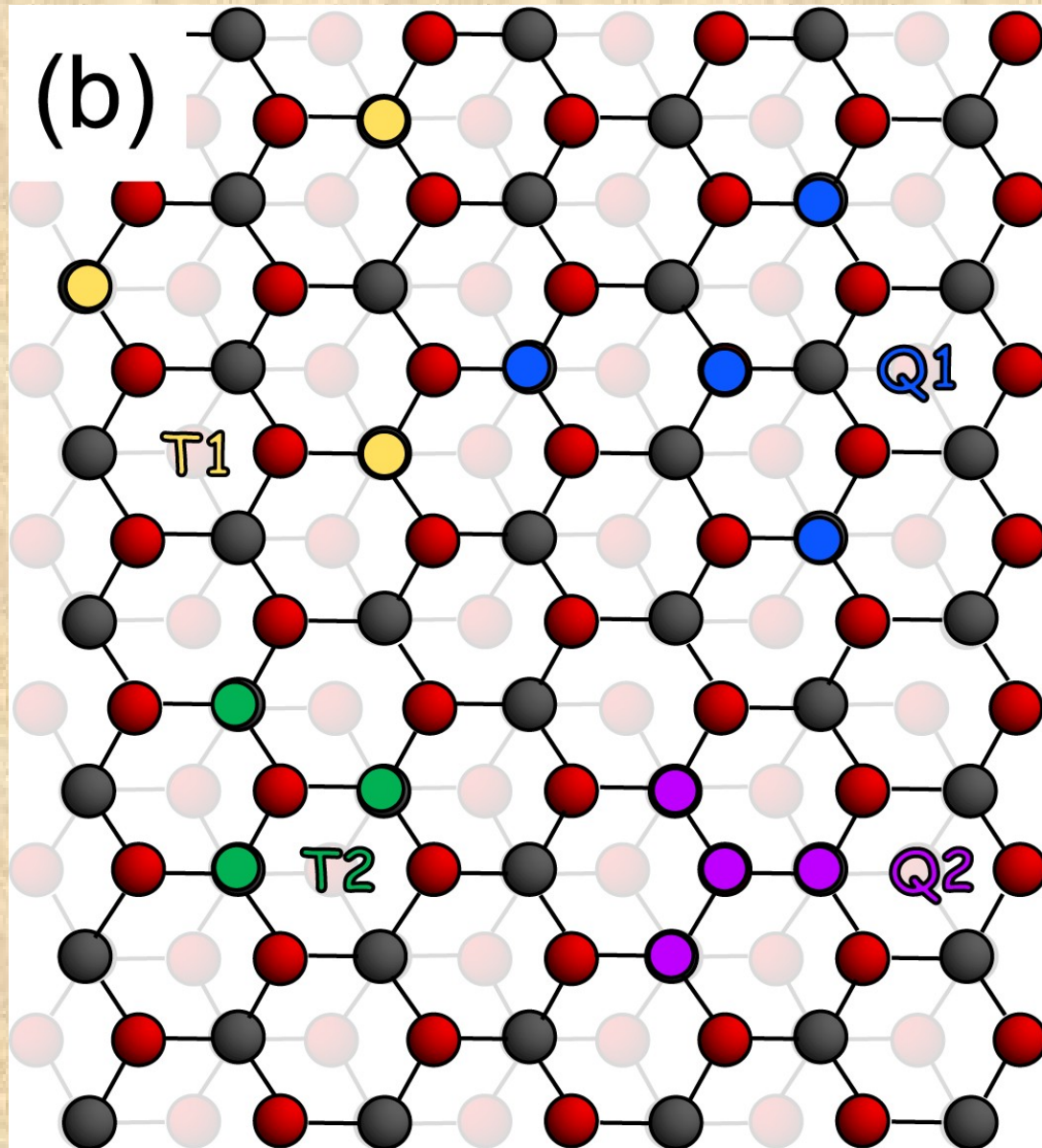


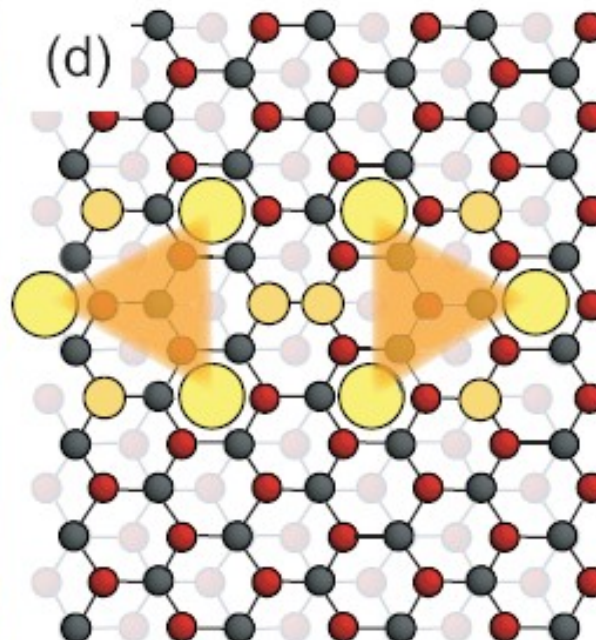
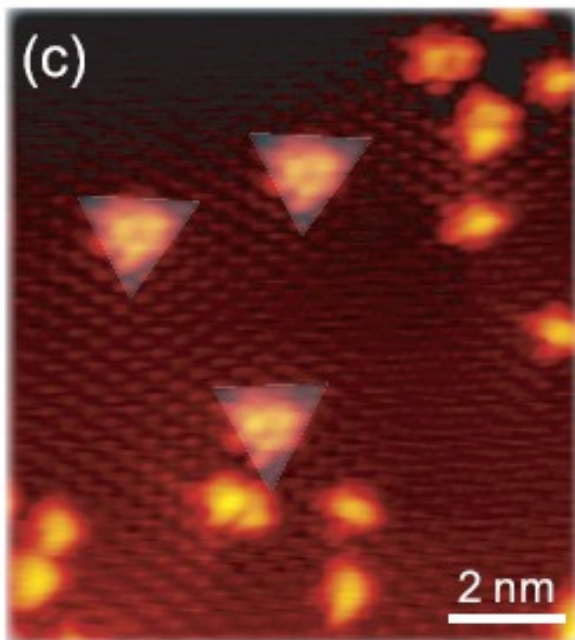
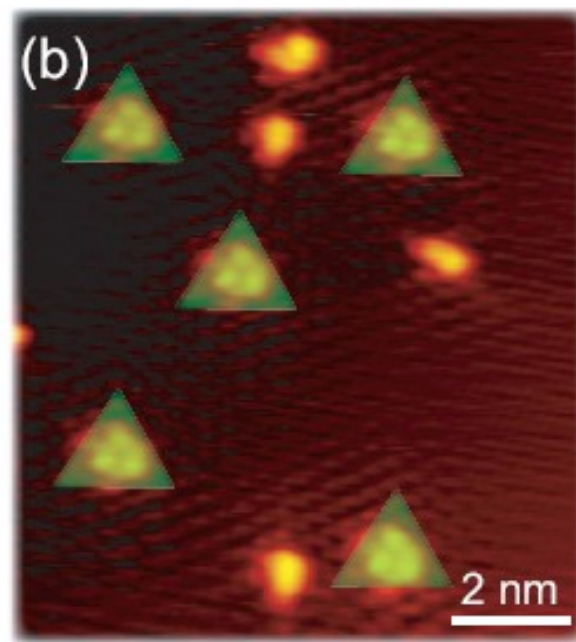
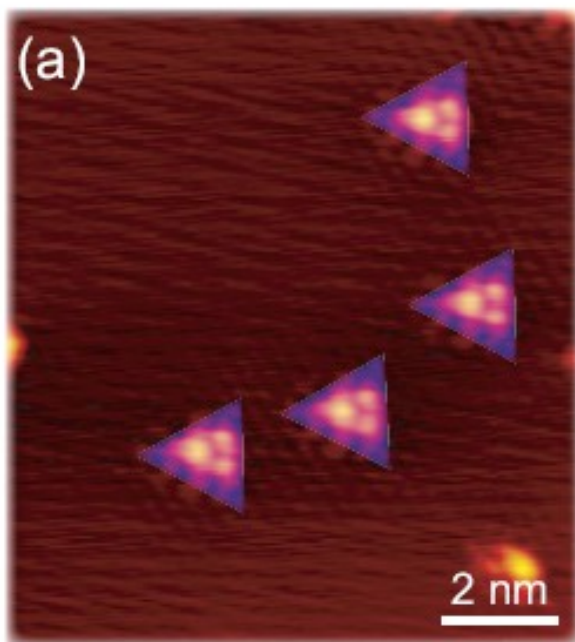
$V_t = -743$ mV, $I_t = -0.62$ nA



Stars – superimposed graphite network

Star structure?





Eagle-Nebula

10-1000 K, 100-10.000 atoms/cm³

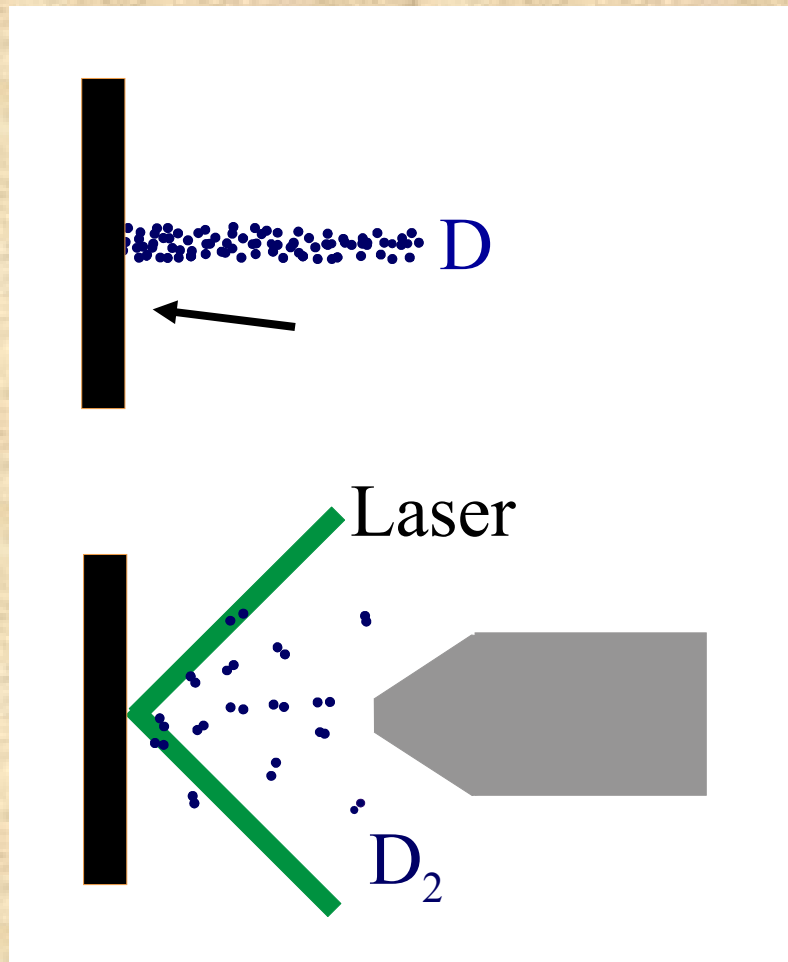


Star-Birth Clouds · M16

HST · WFPC2

PRC95-44b · ST ScI OPO · November 2, 1995
J. Hester and P. Scowen (AZ State Univ.), NASA

Measuring the kinetic energy of formed molecules



Laser Induced
Thermal Desorption
(LITD)

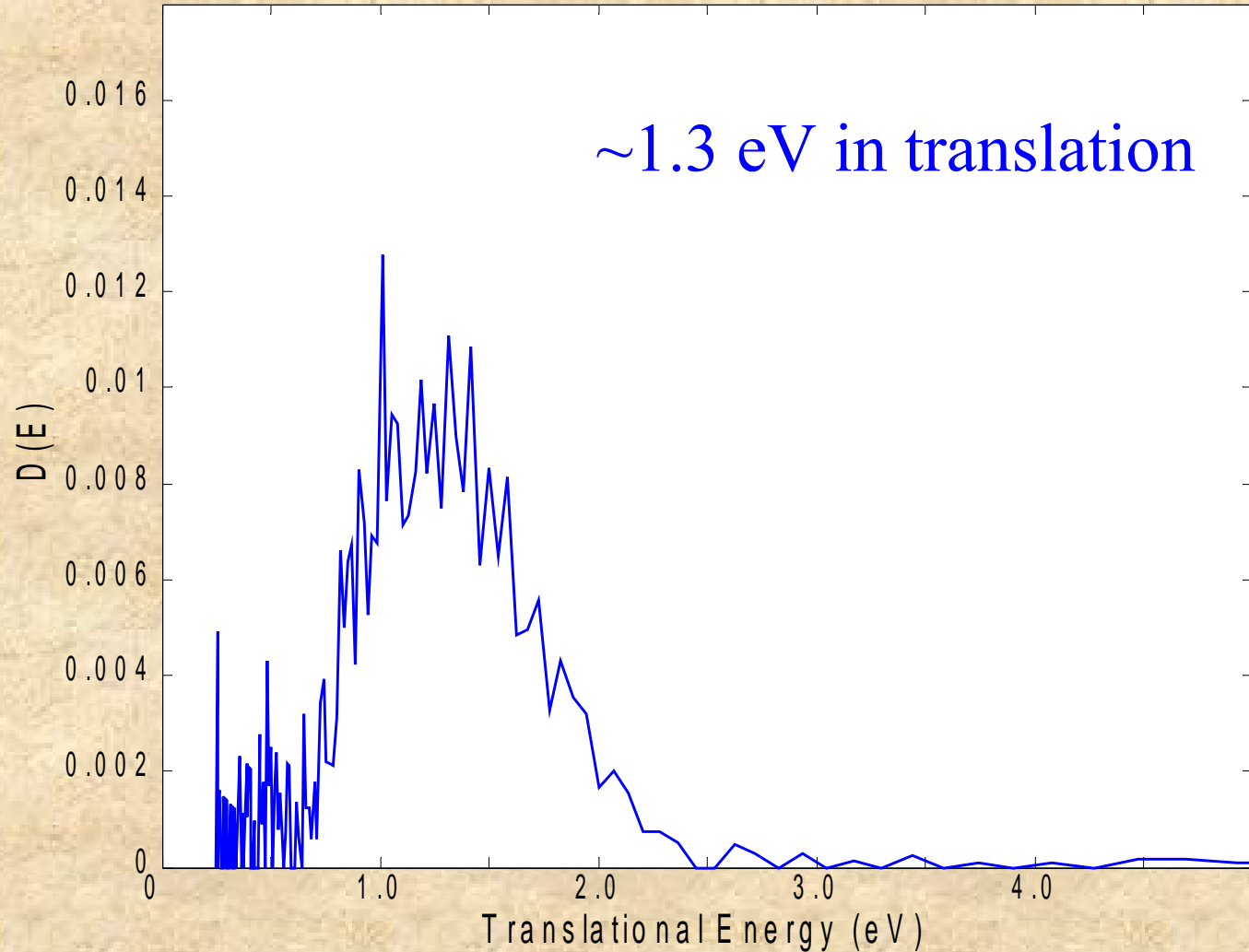
Alexandrite
Laser

4 mJ
100 ns pulse

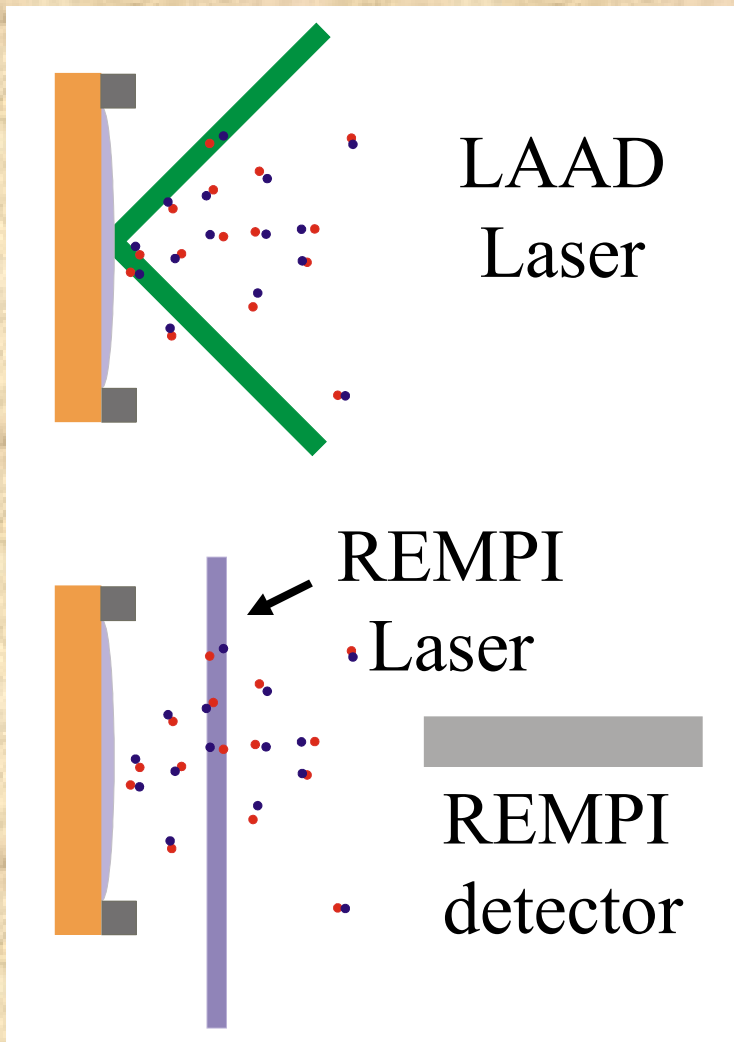


Time of Flight Measurement

Kinetic energy of D_2 formed on graphite



Measuring internal state distributions



3.4 eV liberated

~1.3 eV in translation

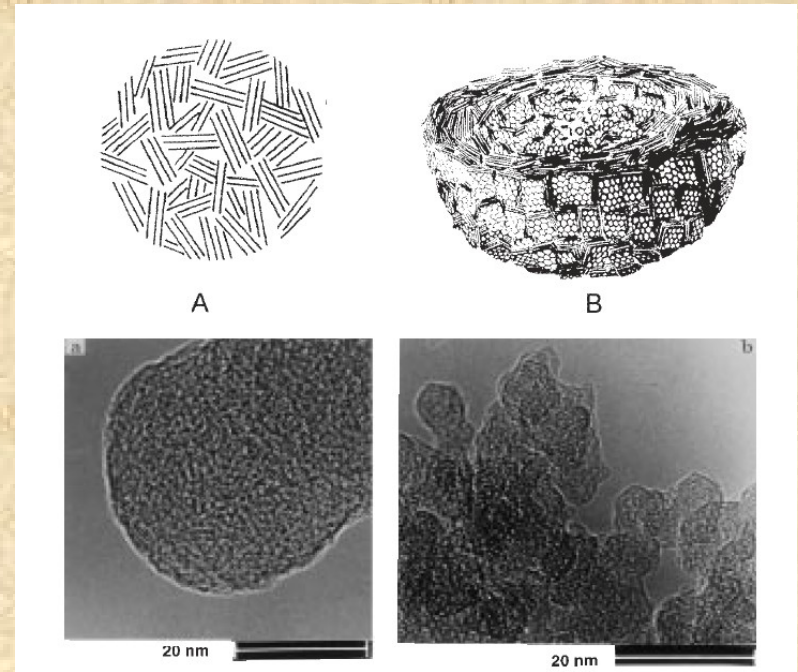
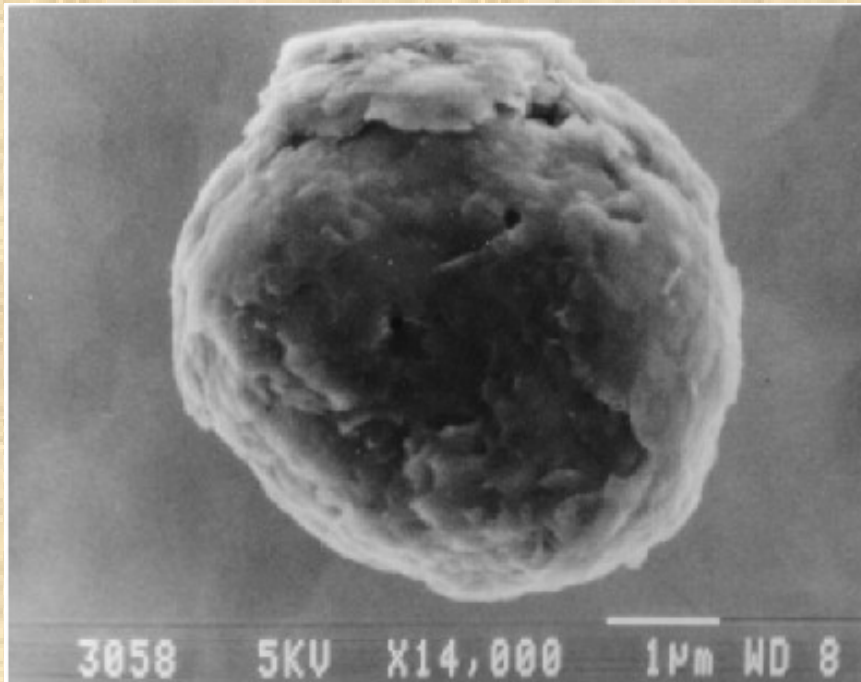
0-1.4 eV in surface heating

0.7 - 2.1 eV in ro-vibration

Talk by Saoud Baouche

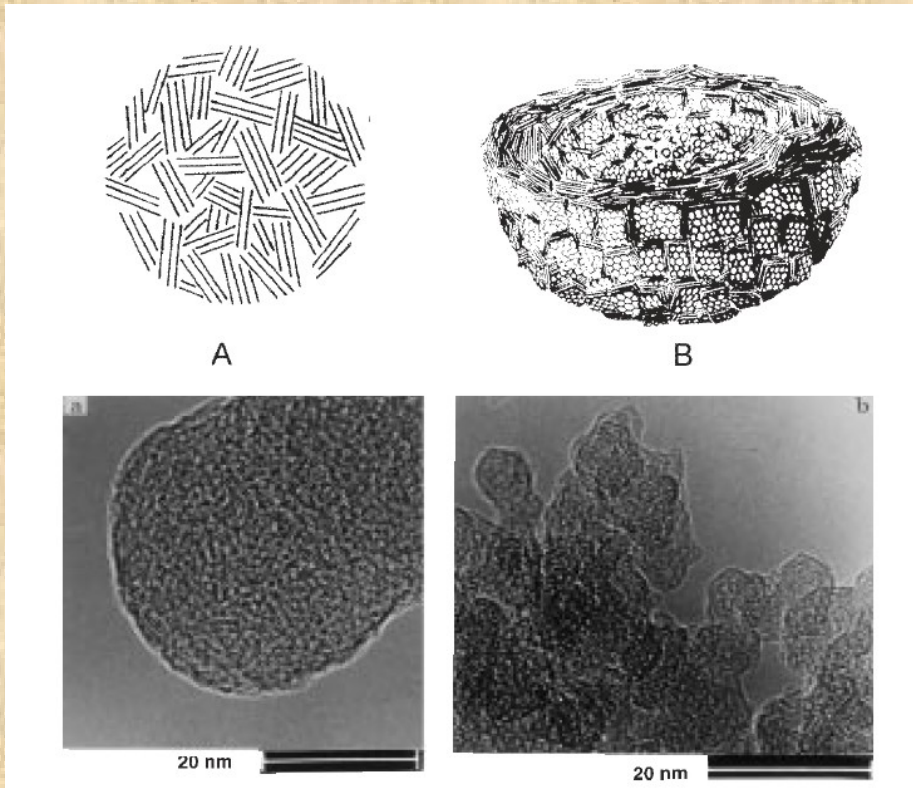
Real carbonaceous surfaces

Onions

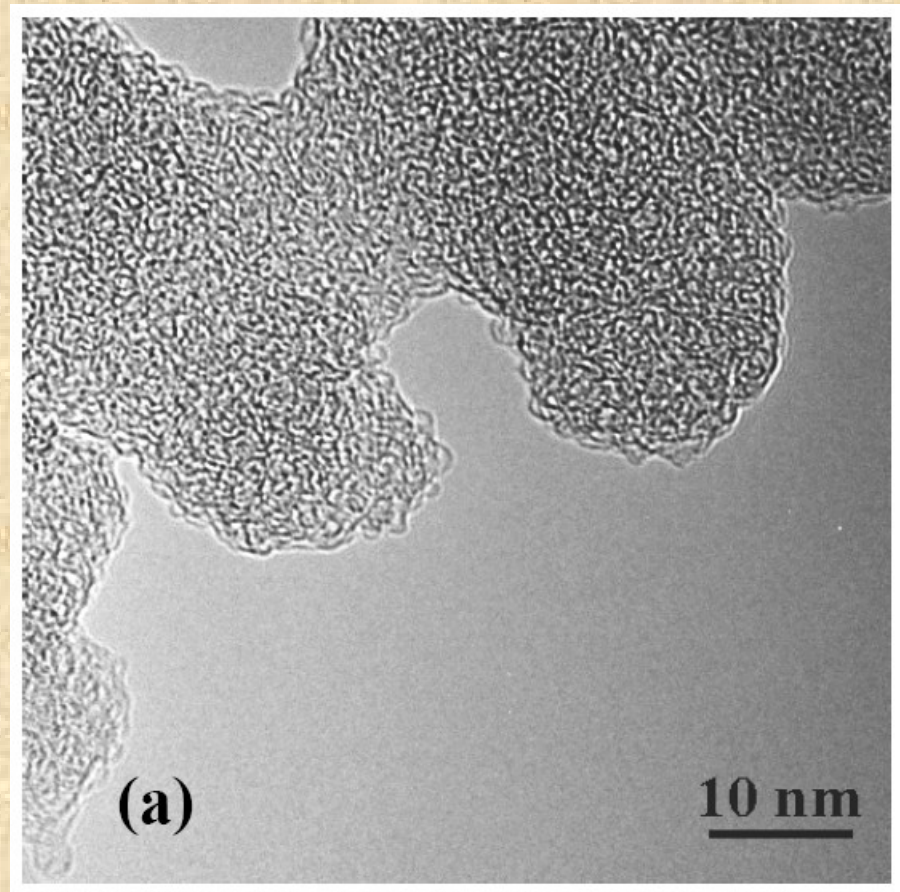


Th. Henning and F. Salama,
Science 282, 2204 (1998)

Porous carbonaceous surfaces



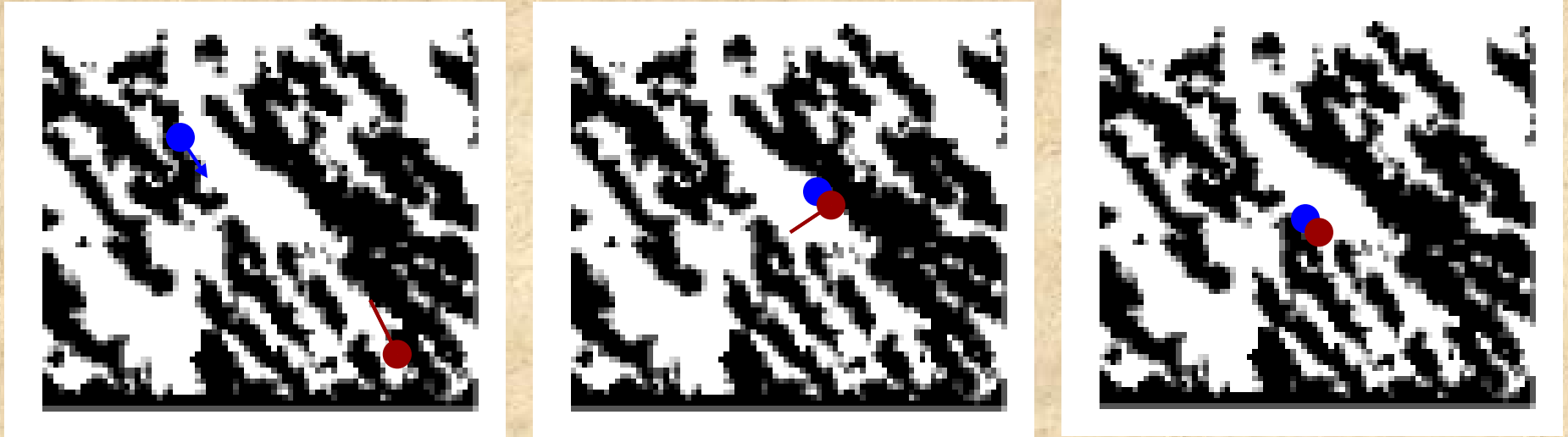
Th. Henning and F. Salama,
Science **282**, 2204 (1998)



Schnaiter et al. Ap. J. **519**, 687 (1999)

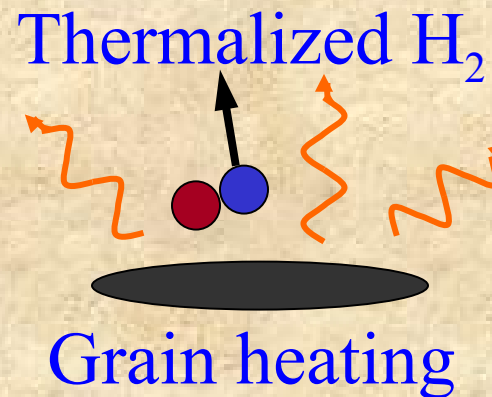
H₂ formation on ASW grown at 10 K:

Physical (Cartoon) picture

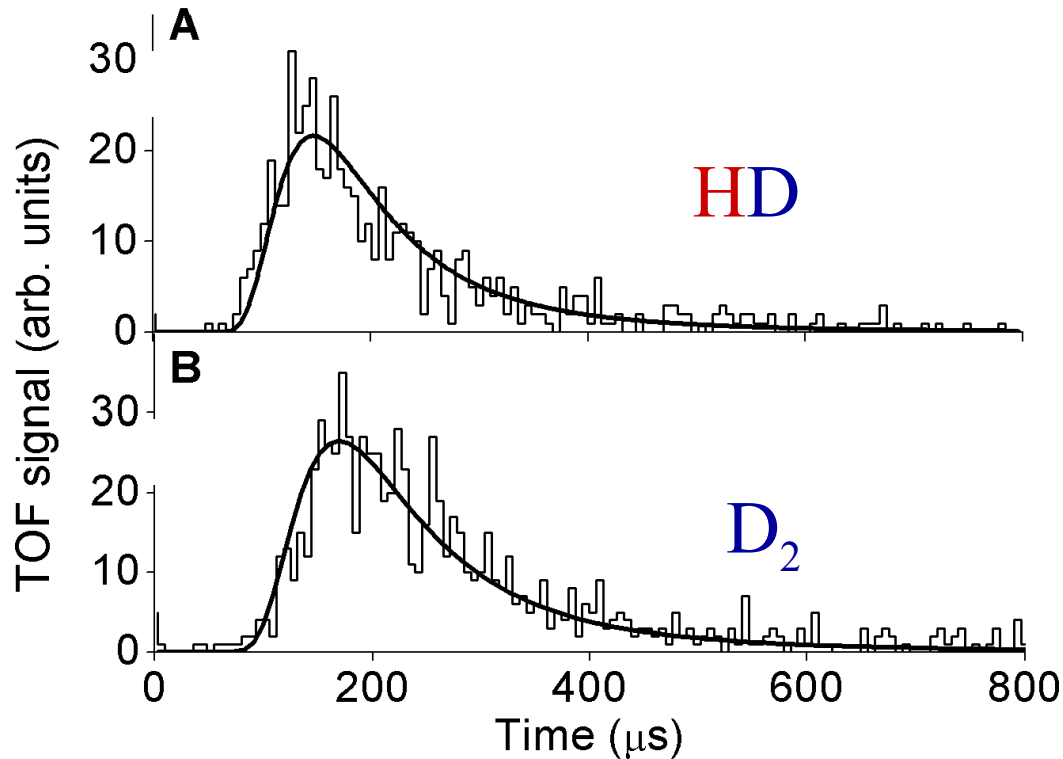


ASW figure from Kimmel et al, JCP 114 p.5295 (2001)

Due to morphology:

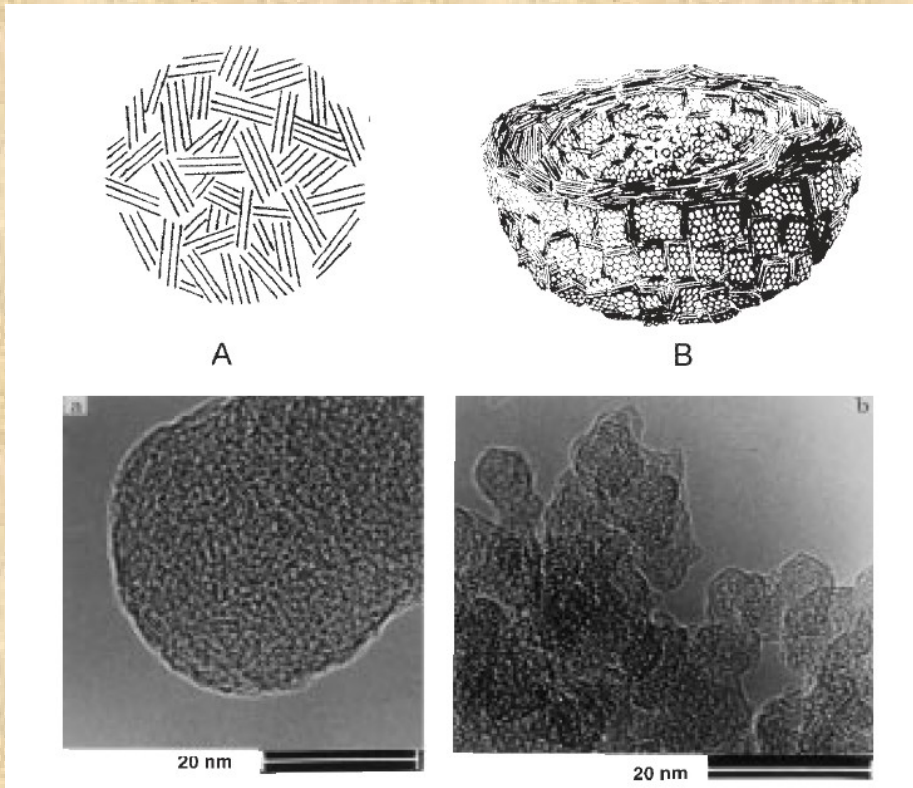


Kinetic energy of HD formed on porous water ice (ASW)

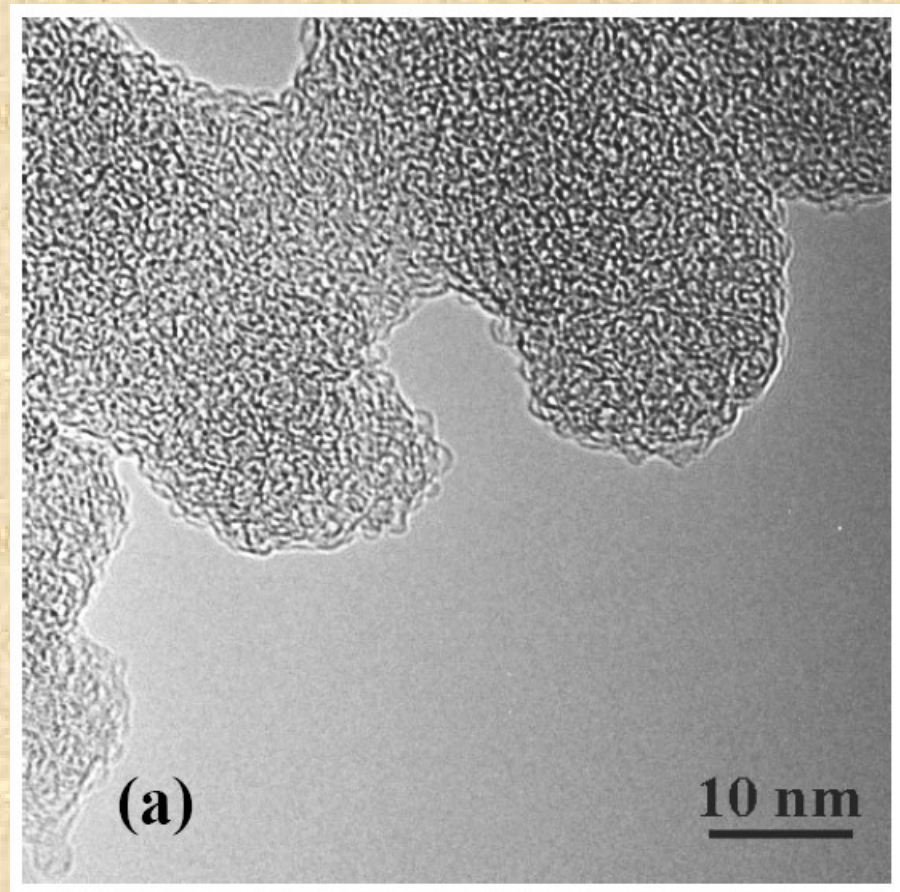


Solid line:
45 K Maxwell-Boltzmann velocity dist.

Porous carbonaceous surfaces

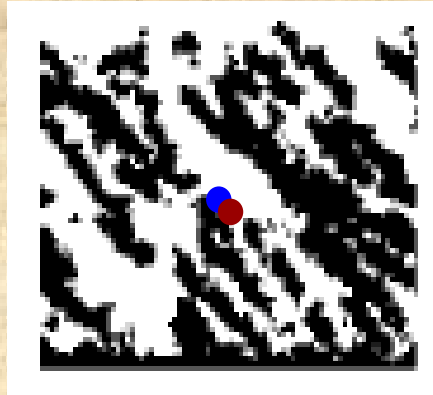
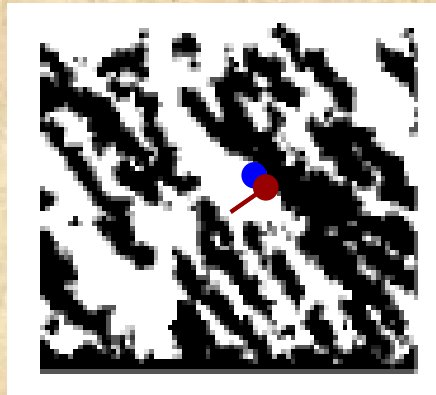
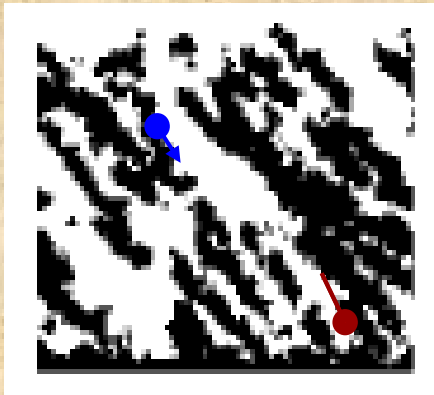


Th. Henning and F. Salama,
Science **282**, 2204 (1998)



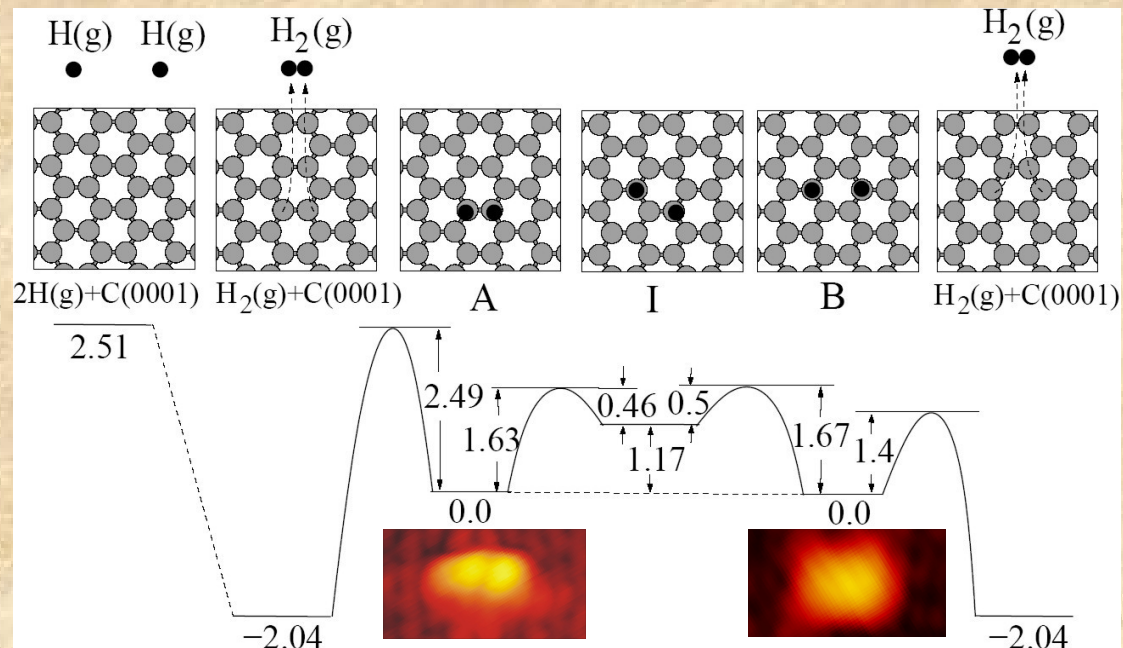
Schnaiter et al. *Ap. J.* **519**, 687 (1999)

Surface astrochemistry: Morphology is a key issue



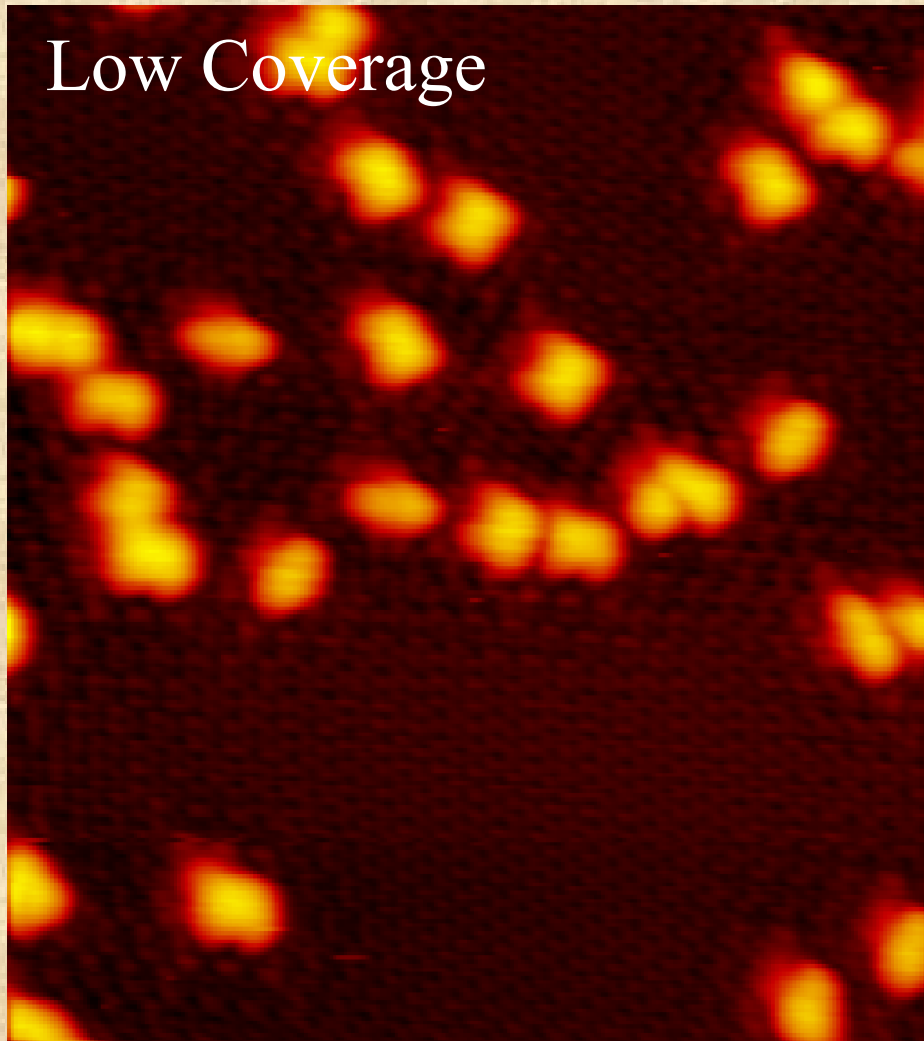
**H₂ formation
on water ice**

**H₂ formation
on graphite**



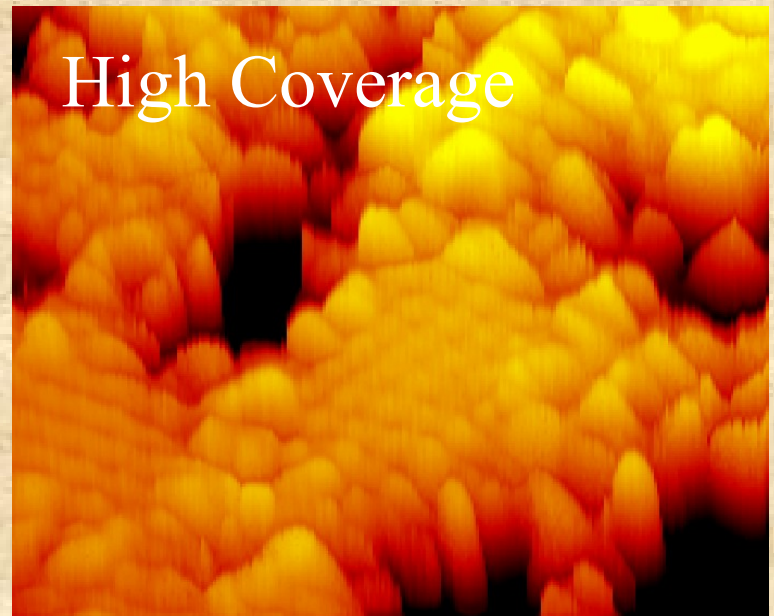
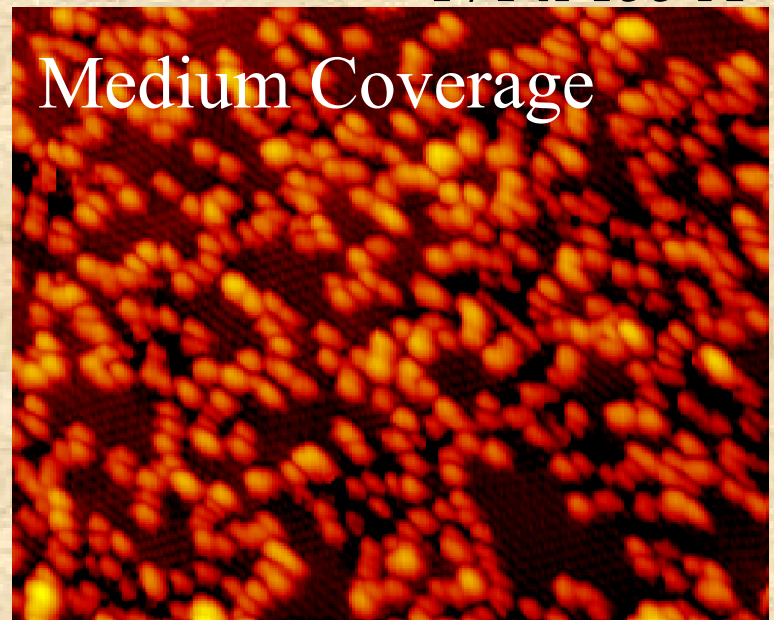
H on HOPG

171 x 155 Å²



103 x 114 Å²

$V_t \sim 800\text{mV}$, $I_t \sim 0.15\text{-}0.2\text{nA}$



80 x 72 Å²

People involved

STM group:

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Flemming Besenbacher

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Dept. Phys. and Astron.

University of Aarhus

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Eva Rauls

Bjørk Hammer

iNANO and

Dept. Phys. and Astron.

University of Aarhus

Zeljko Sljivancanin

EPFL

Laser desorption:

Saoud Baouche

Arnd Baurichter

Victor Petrunin

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and MPG Plasma