



## Role of the surface relaxation.



NGC 346 - Small Cloud of Magellan -

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1. Introduction The Interstellar Medium (ISM)



## Why are we interested in H<sub>2</sub> in the ISM?



Rovibrational state of H<sub>2</sub> vibrationally cold v = 0.5 \*rotationally warm j = 0-30 \*\*

\* T. Giannini et al. Astron. Astrophys. **419**, 999 (2004) \*\* D. Rosenthal et al. A&A **356**, 705 (2000)

**PAH** (*Polycyclic Aromatic Hydrocarbon*) Small Grains (Carbonaceaous) **Big Grains (Silicate)** 

Dust: grains (1%)



Interplanetary /Comet Grains

## **Diffuse clouds**

## **PhotoDissociation Regions**



Dissociation of H<sub>2</sub> very efficient (UV) H>>H, abondance of H<sub>2</sub>?







**Highly efficient recombination Mechanism of H atoms** 

#### **1.** Introduction Main mechanisms on surface









→ Crucial to understand the distribution of the released energy

F

#### 1. Introduction **Previous Works**

Potential: V. Sidis, L. Jeloaica, A.G. Borisov and S.A. Deutscher dans "Molecular hydrogen in space", (Cambridge University Press2000 pp.89-97.)



**1. Introduction** Previous Works

Vibrational distribution of H<sub>2</sub>







Good agreement between Quantum and Quasi-classical dynamics



Form  $H_2$  in lower rovibrational state as observed in the ISM Take into account the full relaxation of the surface

Study of the formation of H<sub>2</sub> on a surface of 200 carbon atoms

606 degrees of freedom



**Classical molecular dynamics** 

**PES of the system?** 

Quantum calculation (DFT) are presently incompatible with molecular dynamics on the fly

Semi-empirical potential Bond-order potential (D.W. Brenner)

**1. Graphene-H-H potential** Construction

Brenner empirical potential \* Used to study small hydrocarbons

\* D.W Brenner Phys.Rev.B. 42, 9458 (1990)

$$E = \sum_{l} E_{l} \quad E_{l_{(ij)}} = \left[ V_{R}(r_{ij}) - \overline{B}_{ij} V_{A}(r_{ij}) \right] \quad \overline{B}_{ij} = \frac{B_{ij} + B_{ji}}{2} + \frac{F_{ij}(xi, xj)}{2}$$



Use of DFT calculations Sidis & Jeloaica

Rougeau & Teillet-Billy

Selection of crucial points for the dynamics

## Chemisorption well

 $z_{c} = 0.36 \text{ Å}$   $z_{H} = 1.49 \text{ Å}$ V = -0.47 eV

## **Chemisorption barrier** :

=	0.13 Å
=	1.86 Å
=	0.26 eV

Corrugation of the surface Jeloaica and Sidis *Chem.Phys.Letters* 1999









Z<sub>C</sub>

 $\mathsf{Z}_{\mathsf{H}}$ 

V

**1. Graphene-H-H potential** Validation



## 1H : Collinear approach of H on a C atom (TOP site)





**Chemisorption Well**:

**1. Graphene-H-H potential** Validation









#### 1. Results **Collinear approach**





**1. Results** Non-collinear approach

Relaxation of all the surface (200 atoms)



$$E_{col} \begin{cases} 0.015 \\ 0.05 \text{ eV} \\ 0.2 \end{cases}$$

>10 000 trajectories Velocity-Verlet algorithm

**1. Results** Non-collinear approach

### **Rovibrational distribution of H**<sub>2</sub>



### **Recession chicage** bability



**1. Results** Non-collinear approach



C

## **1. Results** Non-collinear approach

## **Sudden approximation**



**1.** Conclusions Conclusions



## a realistic potential to study the graphene-H-H system

Potential gave rise to the extended puckering

# Cross section behavior is related to the caracteristics of the potential

Significant amount of energy goes into the surface (≈25% of the released energy)

H<sub>2</sub> is formed in lower rovibrational states that found with the constrained relaxation





## Structure et dynamique des systèmes collisionnels

Muriel Sizun Dominique Teillet-Billy Nathalie Rougeau François Aguillon Victor Sidis

## ANR IRONHI Lerma/Lamap Cergy-Pontoise University Lemaire et al.

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